## The observable (ignoring luminosity-dependent bias)



## Systematics in Galaxy Clustering

Ground-based observator	ries CC	COMMON	
Fiber collision Imaging (angular) Instrumental sys. *completeness harder for ground-based	Milky Way Extinction Astronomical foregrounds ( <i>stars, zodiacal light,</i> <i>nearby galaxies</i> ) Redshift accuracy	Reconstruction Estimator Model validity, scale cuts Window estimation & convolution Wide angle effect Covariances (analytic v. mocks) Super-sample covariance ? Noisy?	Interlopers Spectral confusion Instrumental sys. *z-accuracy harder for space-based
		Likelihood shape Projection effect Prior volume, Prior weight HOD prior	

- What are the scales affected: small scales, BAO scale, fNL scales?
- Limiting science measurements for Stage IV, Stage V? i.e. DESI, DESI-II, Euclid, SPHEREX, SPEC-S5, WST
- What are the hurdles to mitigate Data/Model/Likelihood systematics?
- Folks that are leading the analyses: what do you need that you do not have from the Theorists ?
- Folks that are on the theory side: what do you feel needs more investigation in the Surveys main analyses ?
- Are there systematics that the community is ignoring, unaware of?

## **Cross correlations, probe combination, tensions**

Cross correlations of probes that have different systematics can help extracting the signal, are we ready to exploit them?

e.g. optical vs radio, clustering vs lensing

Probe combination needs control of systematics and cross covariances, are we ready for the "cross correlation feast"?

Uncontrolled systematics will lead to tensions -

ready to blind yourself? is blinding the way to go?