

# Evolution mapping: a tool to describe non-linear density fluctuations

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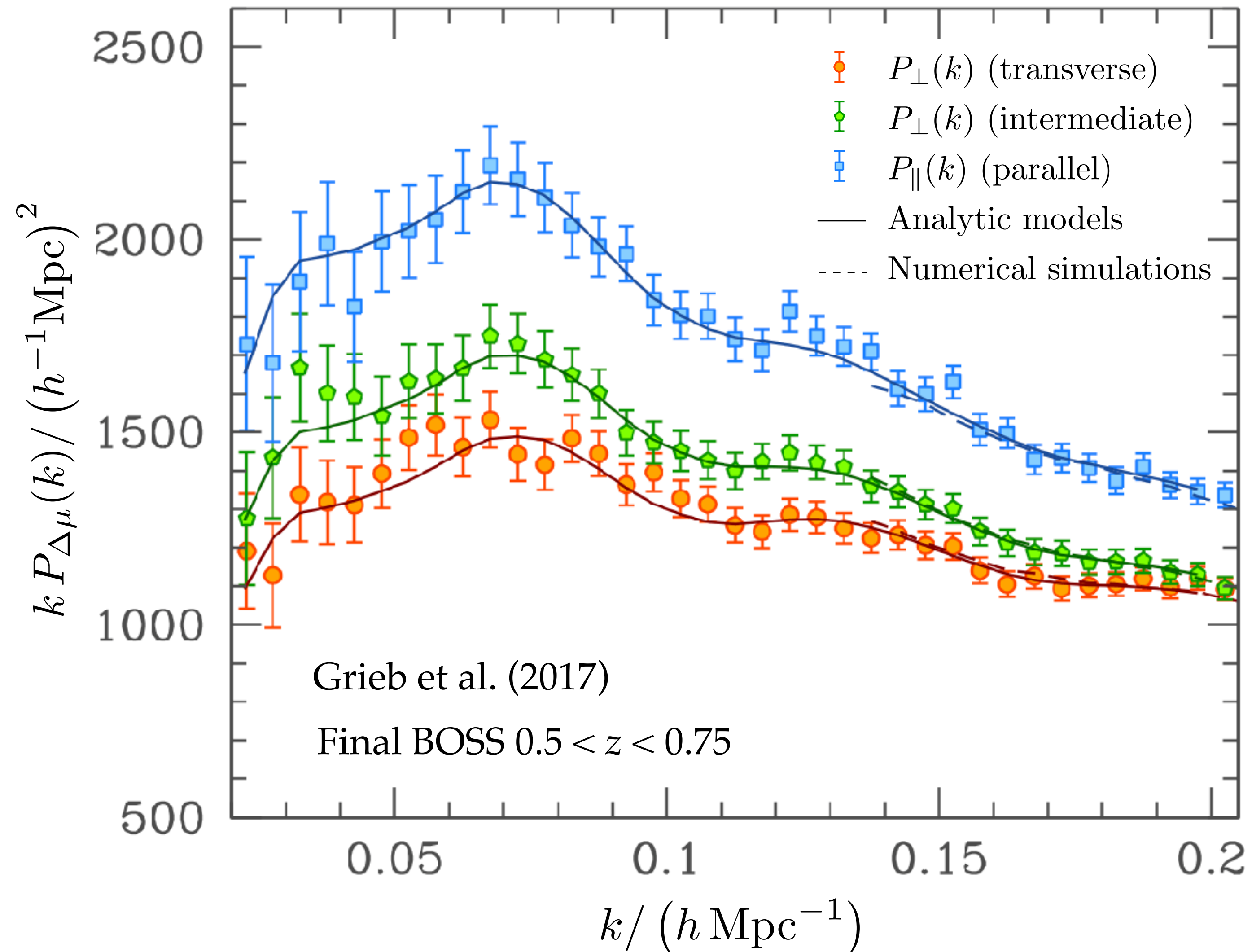
C. Correa, A. Eggemeier, M. Esposito, L. Finkbeiner, A. Fiorilli,  
A. Perez Fernandez, A. Pezzotta, A. Ruiz, A. Semenaite

New Strategies for Extracting Cosmology from Galaxy Surveys II - July 2024

# New LSS analysis methods

State-of-the-art models only access large-scale information.

New surveys exacerbate this problem.



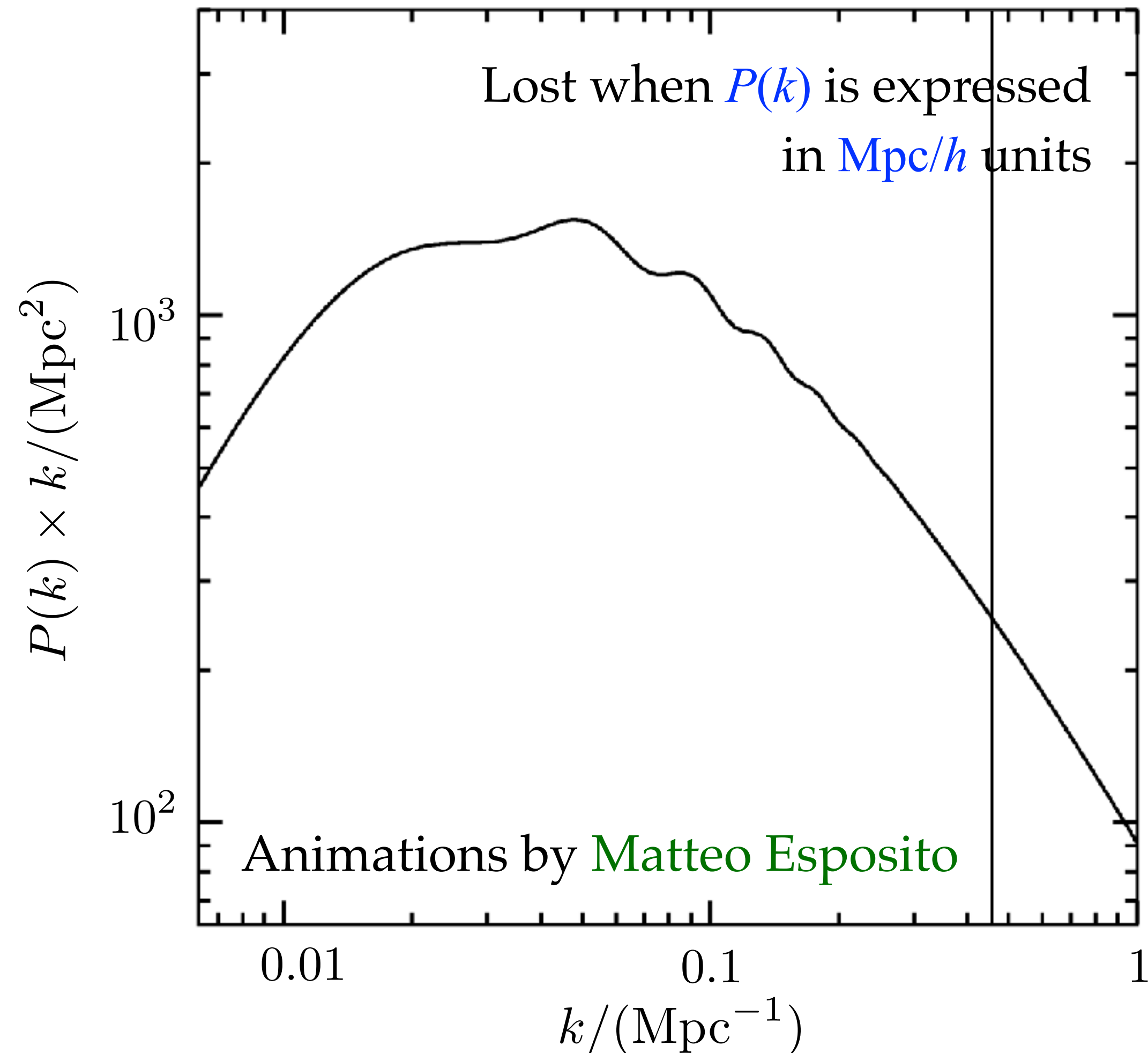
Model predictions based on **simulations**:  
**Emulators** trained on simulations.

**Simulation-based inference.**

**Field-level inference.**

# Evolution mapping: linear $P(k)$

We can classify cosmological parameters according to their impact on  $P(k)$



shape evolution

$$\Theta = (\omega_c, \omega_b, n_s, A_s, \omega_K, \omega_{\text{DE}}, w_0, w_a, \dots, z)$$

$$\Theta = (\omega_c, \omega_b, n_s, \sigma_{12})$$

~~$$h^2 = \sum_i \omega_i,$$~~

~~$$\Omega_i = \omega_i / h^2,$$~~

~~$$\sigma_{8/h} = \sigma(R = 8/h \text{ Mpc})$$~~

$$\sigma_{12} = \sigma(R = 12 \text{ Mpc})$$

$$P_L(k|z, \Theta_s, \Theta_e) = P_L(k|\Theta_s, \sigma_{12}(z, \Theta_s, \Theta_e))$$

# Alternative normalizations

The scale  $R = 12 \text{ Mpc}$  is arbitrary.

Alternative normalizations include:

- The value  $\sigma(R)$  at any scale defined in  $\text{Mpc}$ .
- The value of  $\Delta^2(k_p)$  at any scale  $k_p$  defined in  $\text{Mpc}^{-1}$ .
- The scale  $R_{n1}$  at which  $\sigma(R_{n1}) = 1$ .
- The scale  $k_{n1}$  at which  $\Delta^2(k_{n1}) = 1$ .

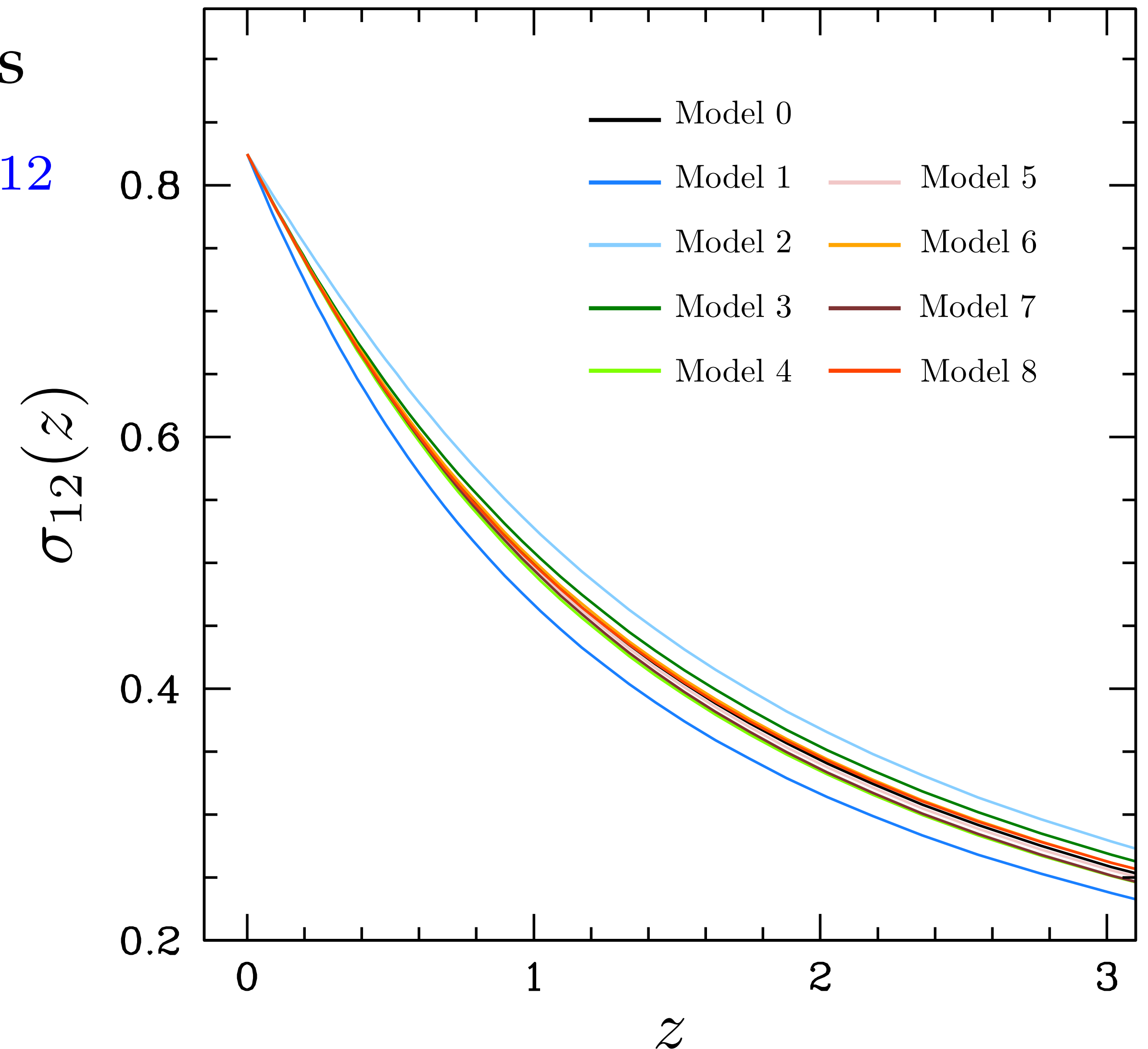
# Evolution mapping: linear theory

We can map the  $z$  evolution of models with identical  $\Theta_s$  using the value of  $\sigma_{12}$

## The Aletheia cosmologies

Model	Definition
Model 0	Reference $\Lambda$ CDM as in Table 1.
Model 1	$\Lambda$ CDM, $\omega_{\text{DE}} = 0.1594$ ( $h = 0.55$ ).
Model 2	$\Lambda$ CDM, $\omega_{\text{DE}} = 0.4811$ ( $h = 0.79$ ).
Model 3	$w$ CDM, $w_{\text{DE}} = -0.85$ .
Model 3	$w$ CDM, $w_{\text{DE}} = -1.15$ .
Model 5	Dynamic dark energy, $w_a = -0.2$ .
Model 6	Dynamic dark energy, $w_a = 0.2$ .
Model 7	Non-flat $\Lambda$ CDM, $\Omega_K = -0.05$ .
Model 8	EDE model, $w_0 = -1.15$ , $\Omega_{\text{DE},e} = 10^{-5}$

*Aletheia*: greek goddess of truth.  
It means “to reveal”.



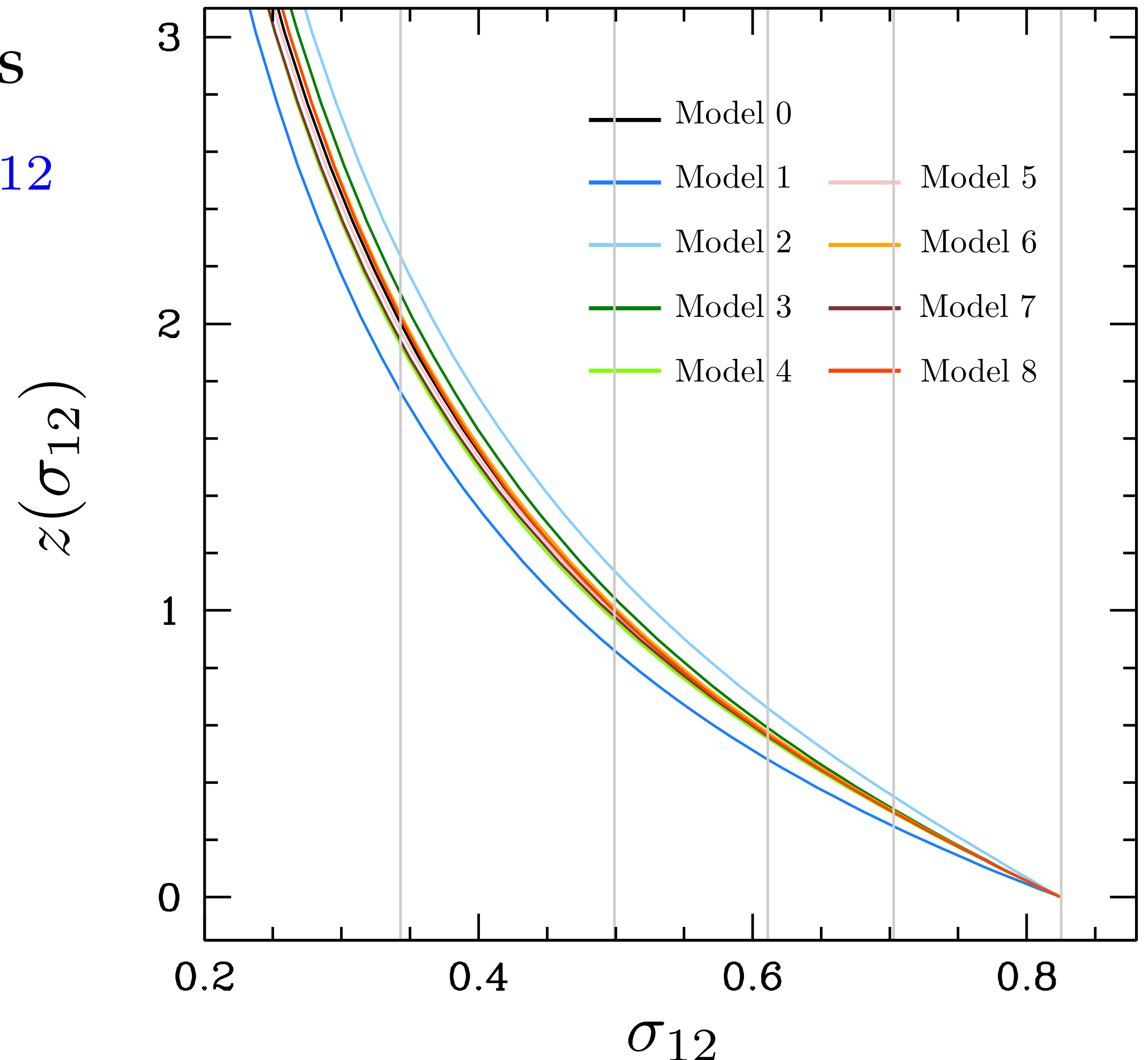
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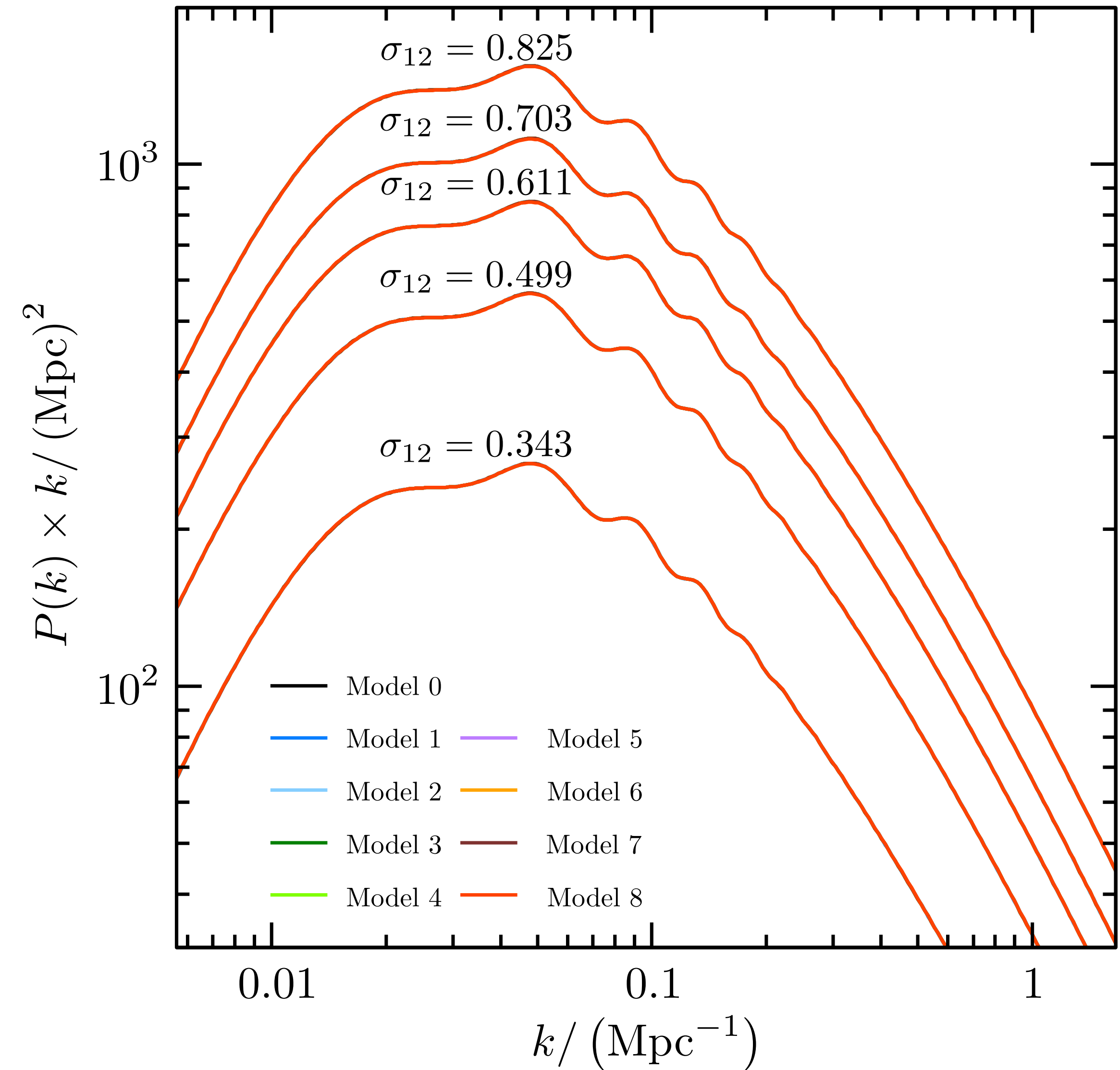
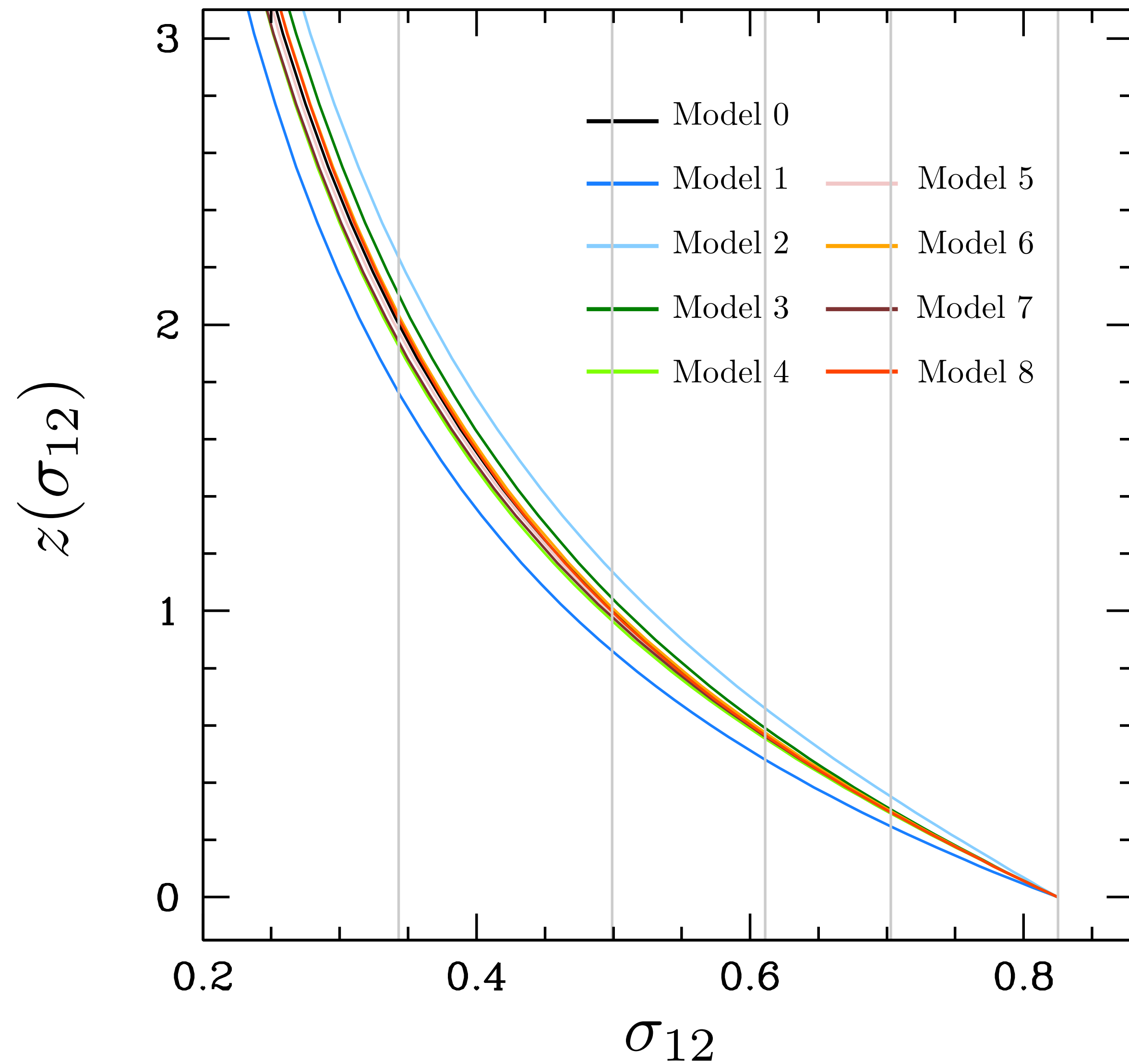
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# Evolution mapping: linear theory

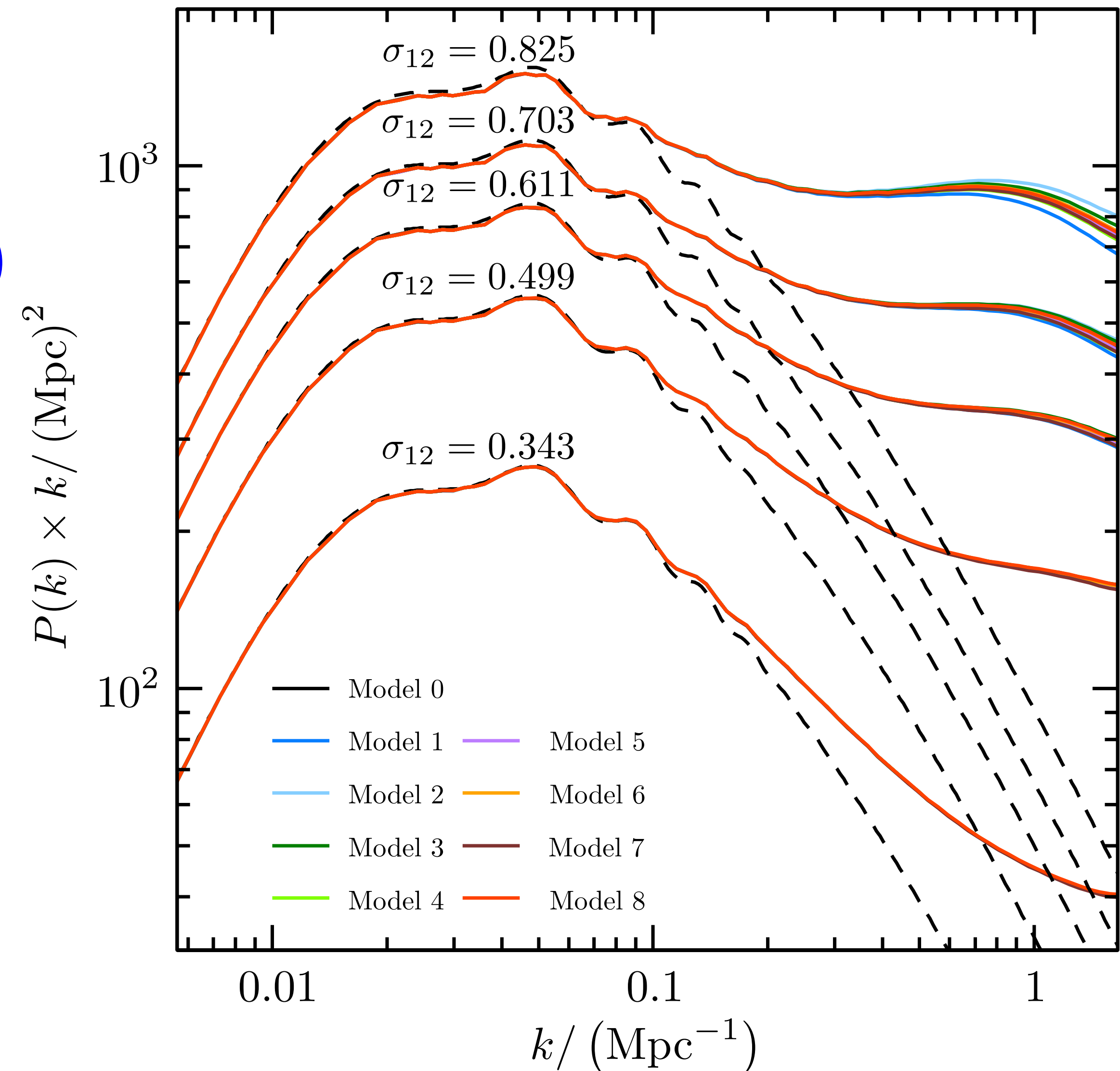


# Evolution mapping: non-linear $P(k)$

- Evolution mapping gives a good description of the non-linear  $P(k)$

$$P(k|z, \Theta_s, \Theta_e) \simeq P(k|\Theta_s, \sigma_{12}(z, \Theta_s, \Theta_e))$$

- Differences can be seen in the deeply non-linear regime.
- Deviations are larger at high  $k$  and increase with  $\sigma_{12}$ .
- The models with the largest deviations change with  $\sigma_{12}$ .



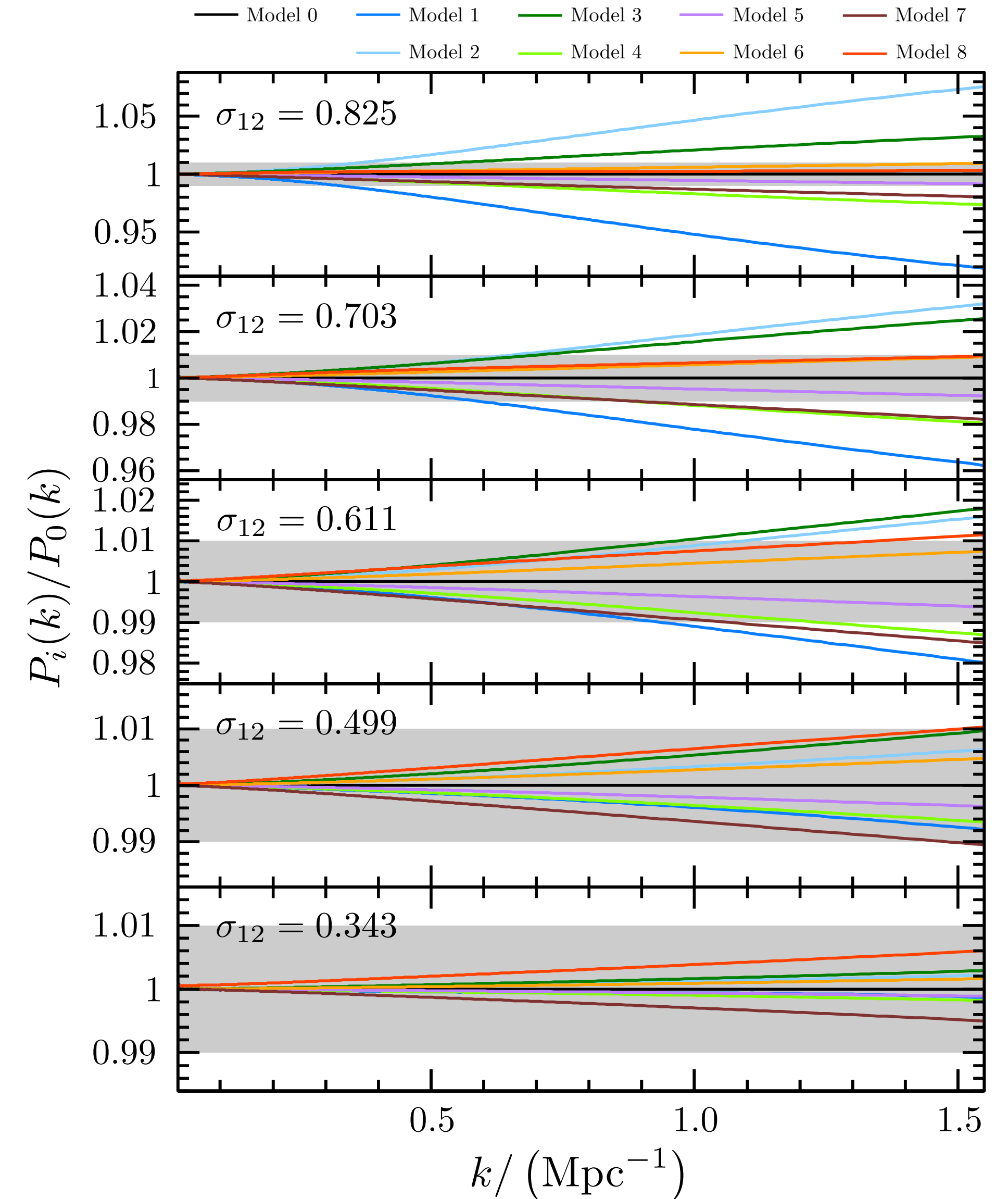


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# Deviations from the $\sigma_{12}$ degeneracy

- The differences could be described as a functional of  $g(a) = D(a)/a$

$$\delta P [\delta g(\sigma_{12})] = \int \frac{\delta P}{\delta g} \delta g(\sigma'_{12}) d\sigma'_{12}$$

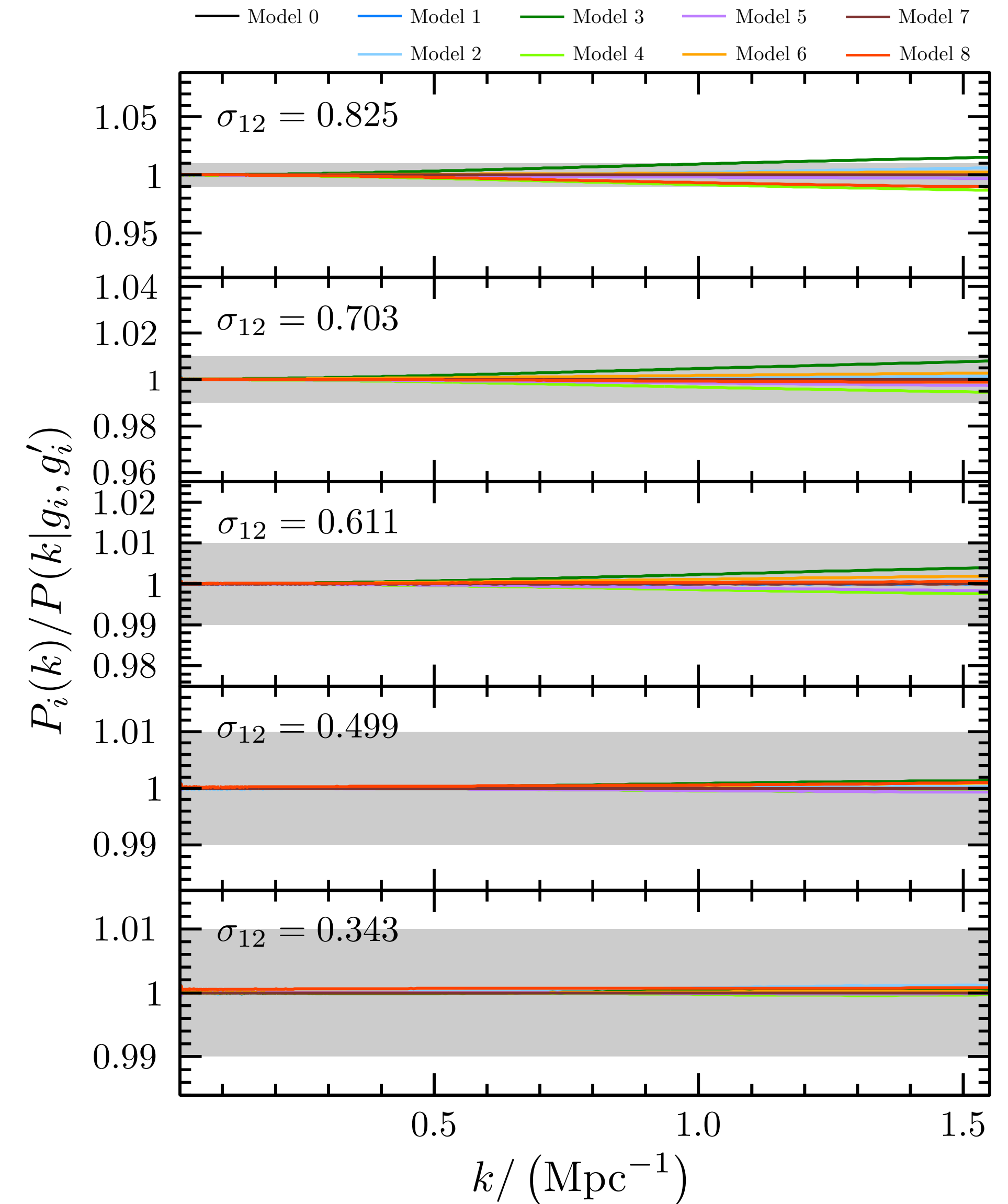
- A simpler recipe:

$$P(k|g, g') = P(k|g_0, g'_0)$$

$$+ \frac{\partial P}{\partial g} (k|g_0, g'_0) (g - g_0)$$

$$+ \frac{\partial P}{\partial g'} (k|g_0, g'_0) (g' - g'_0)$$

$$g' = \frac{\partial g}{\partial \ln \sigma_{12}}$$



# The peculiar velocity field

- Modelling peculiar velocities is essential to analyse redshift-space quantities.
- At the linear level,  $\mathbf{v}$  and  $\delta$  are linked through the continuity equation.

$$\theta := -\frac{\nabla \cdot \mathbf{v}}{af(a)H(a)} \longrightarrow P_{\theta\theta}(k) = P_{\theta\delta}(k) = P_{\delta\delta}(k)$$

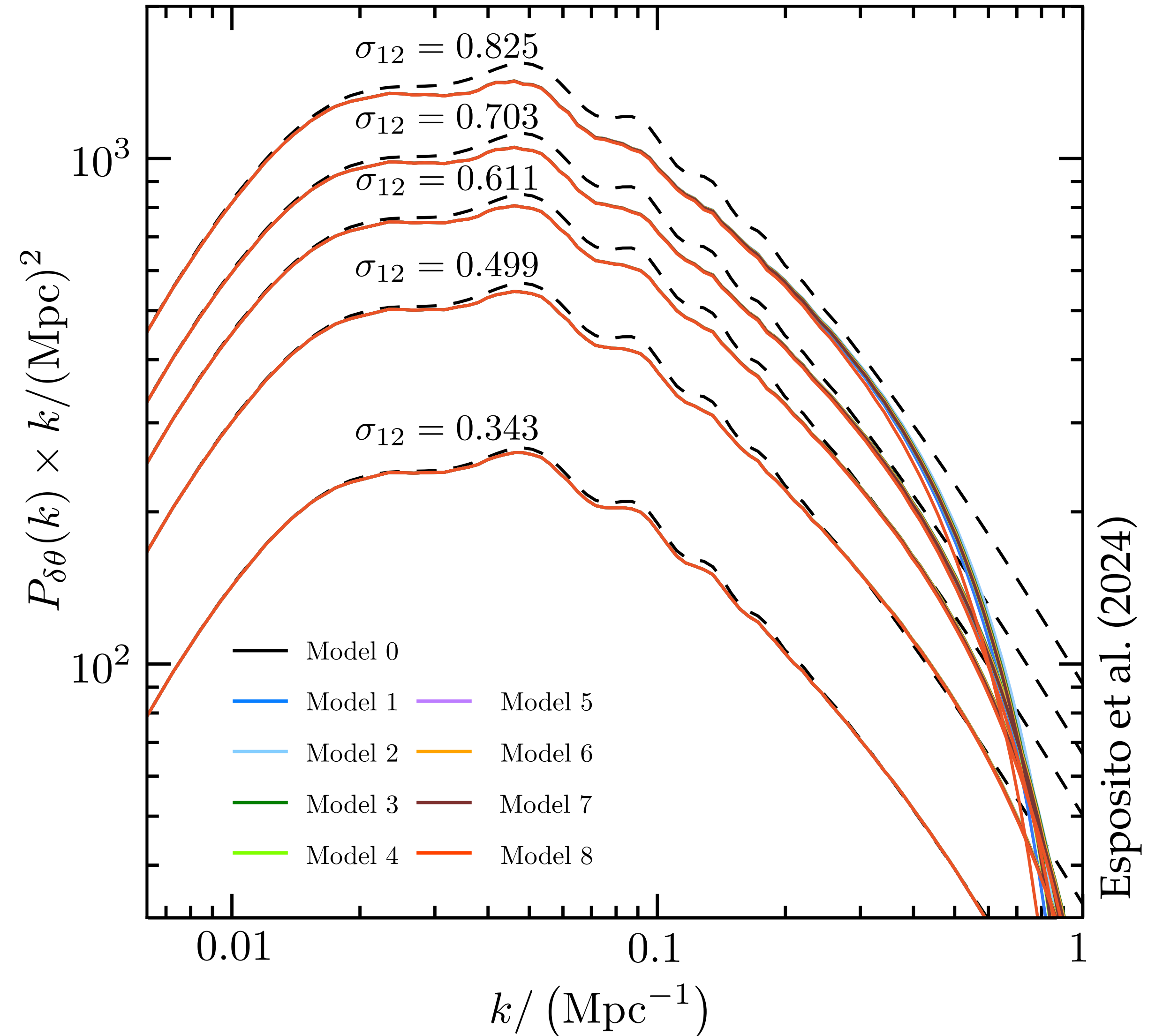
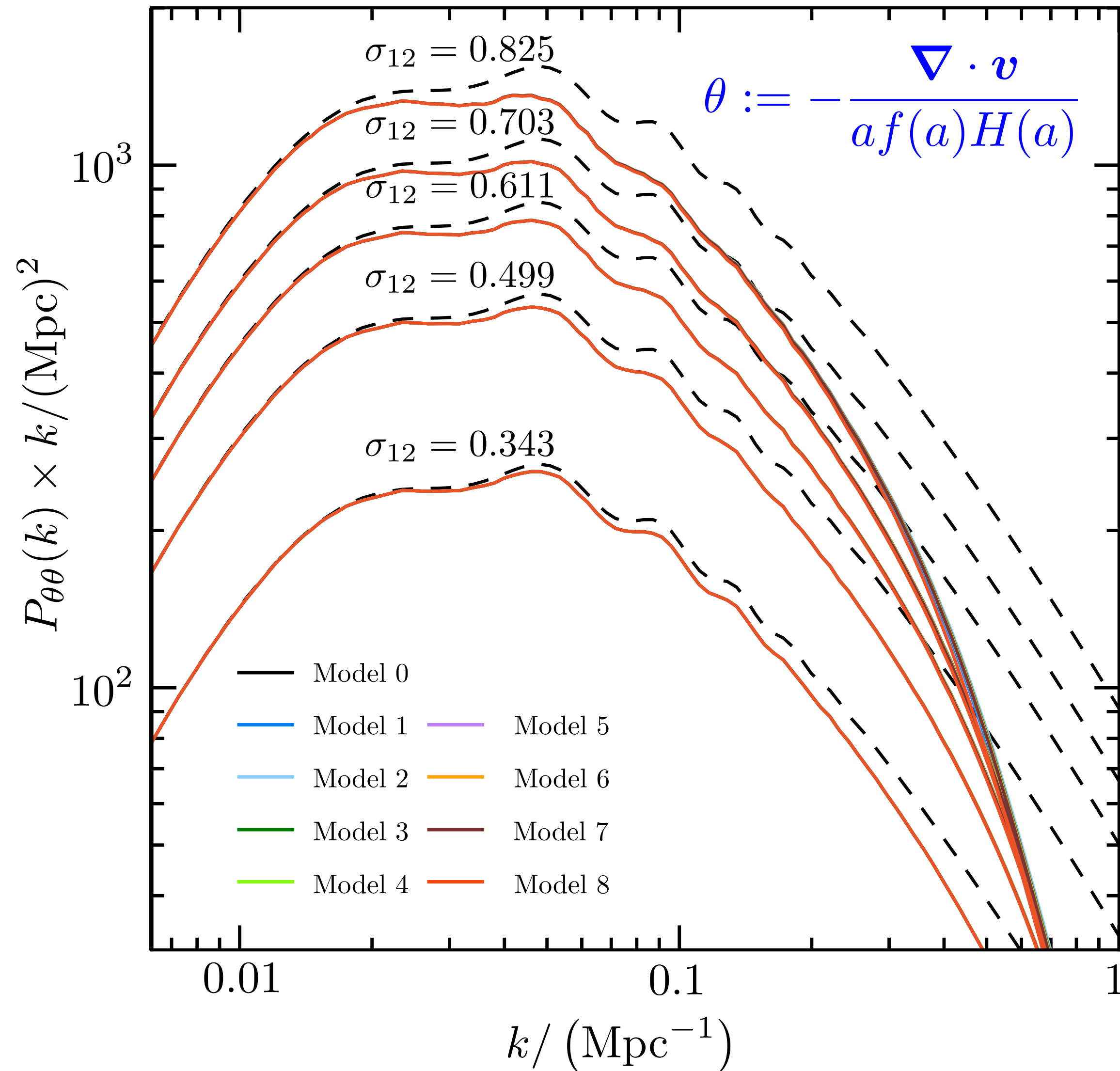
- Considering  $\ln \sigma_{12}$  as a time variable

$$\boldsymbol{\Upsilon} = \frac{d\mathbf{x}}{d \ln \sigma_{12}} = \frac{d\mathbf{x}}{dt} \frac{dt}{d \ln \sigma_{12}} = \frac{\mathbf{v}}{af(a)H(a)}$$

- The rescaled velocities  $\boldsymbol{\Upsilon}$  follow the evolution mapping relation.

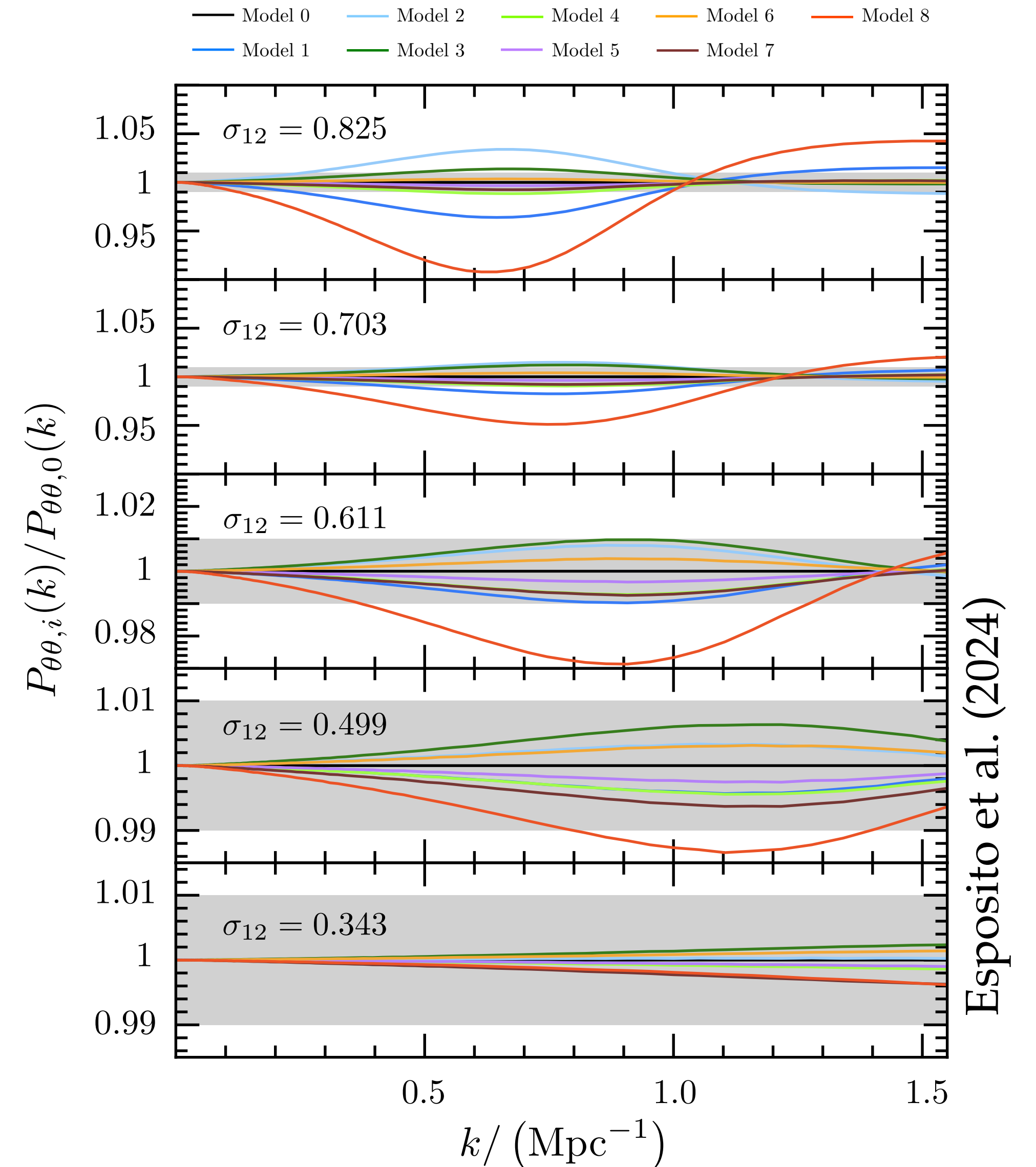
$$\theta = -\nabla \cdot \boldsymbol{\Upsilon}$$

# The peculiar velocity field



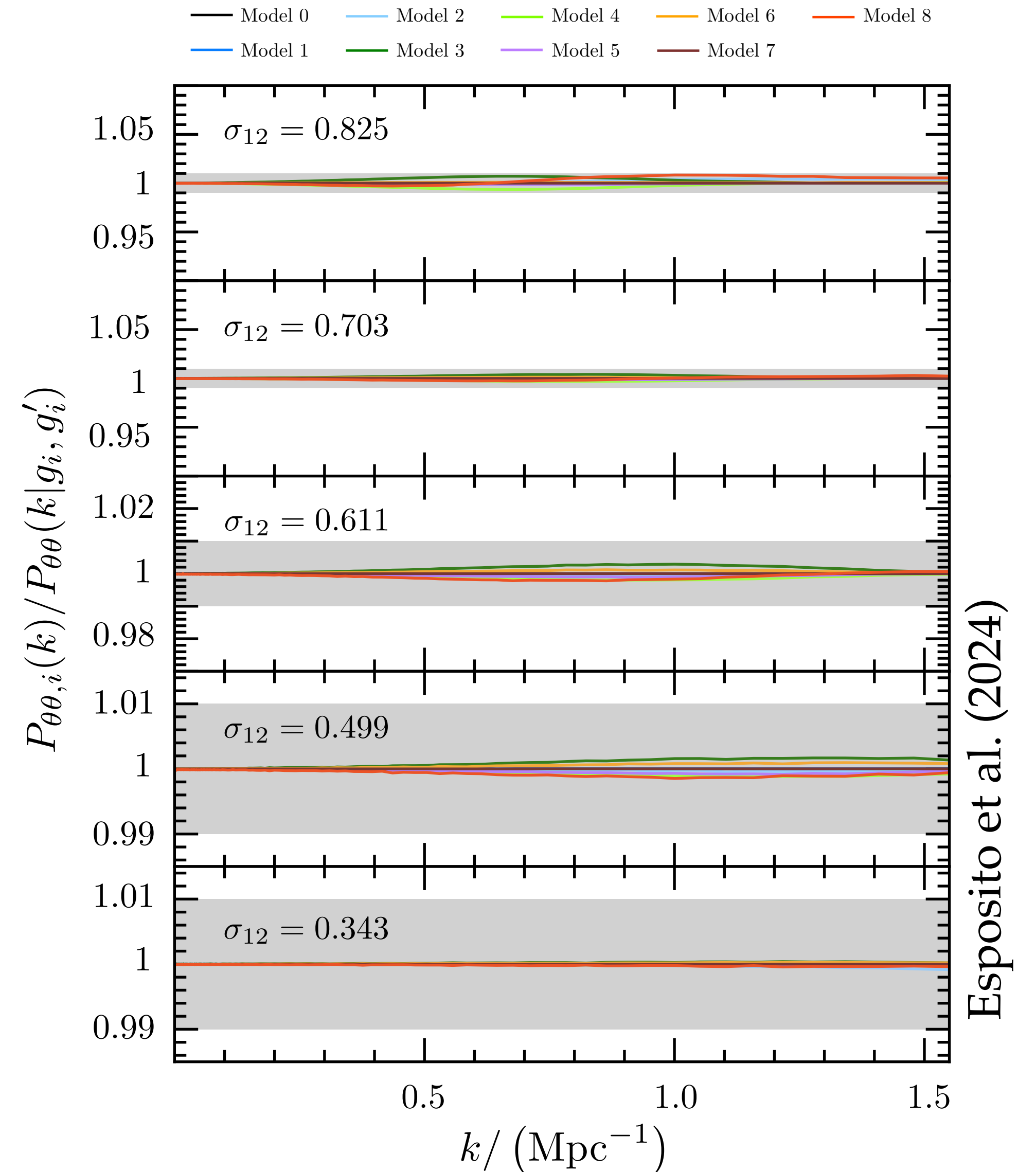
# The peculiar velocity field

- Deviations from a perfect degeneracy follow a similar pattern as  $P_{\delta\delta}(k)$ .
- With the appearance of vorticity, the trend in the deviations is reverted.
- For each model, the maximum deviations are smaller than for  $P_{\delta\delta}(k)$ .
- The differences can also be described in terms of  $\Delta g(\sigma_{12})$  and  $\Delta g'(\sigma_{12})$



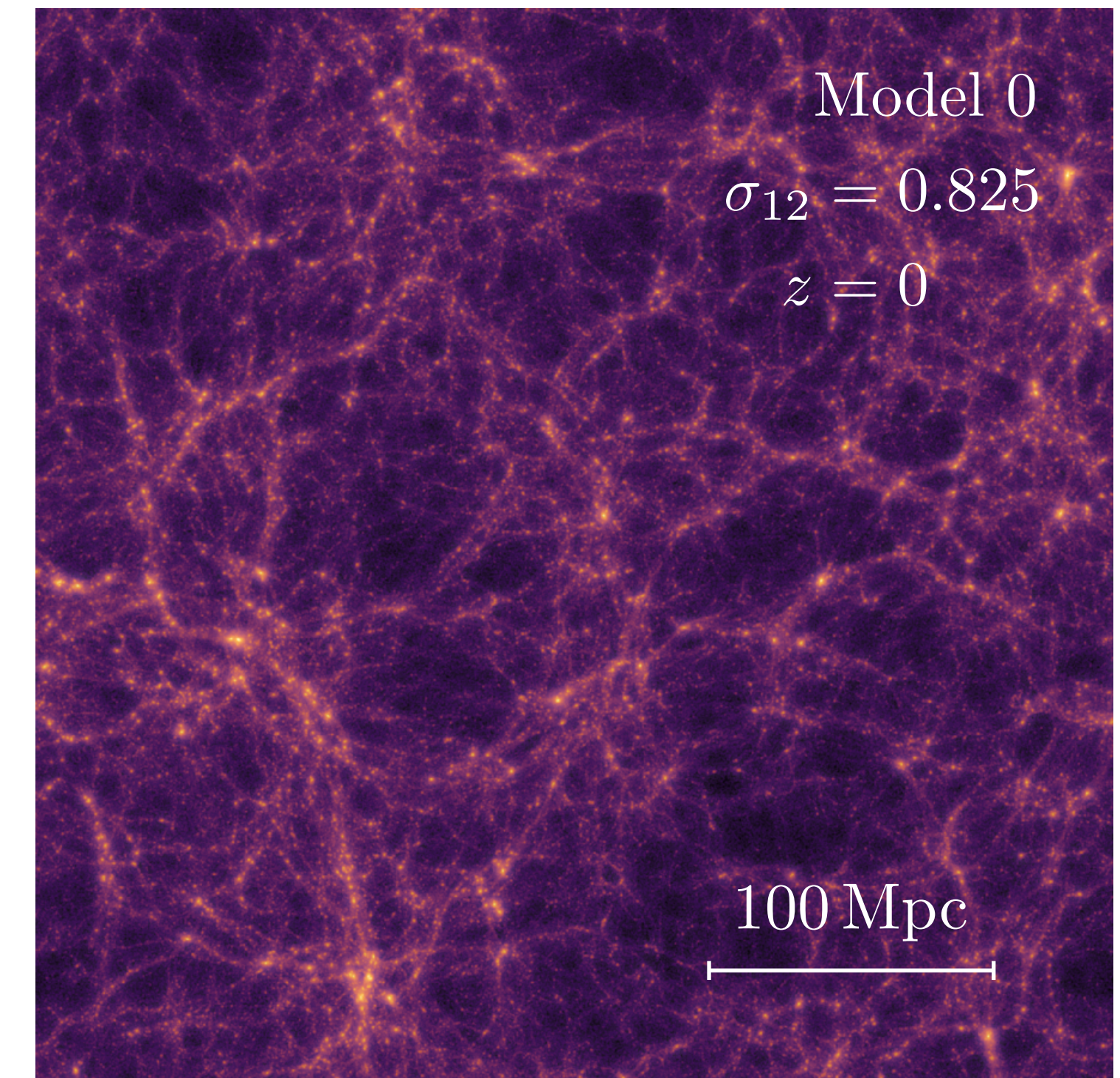
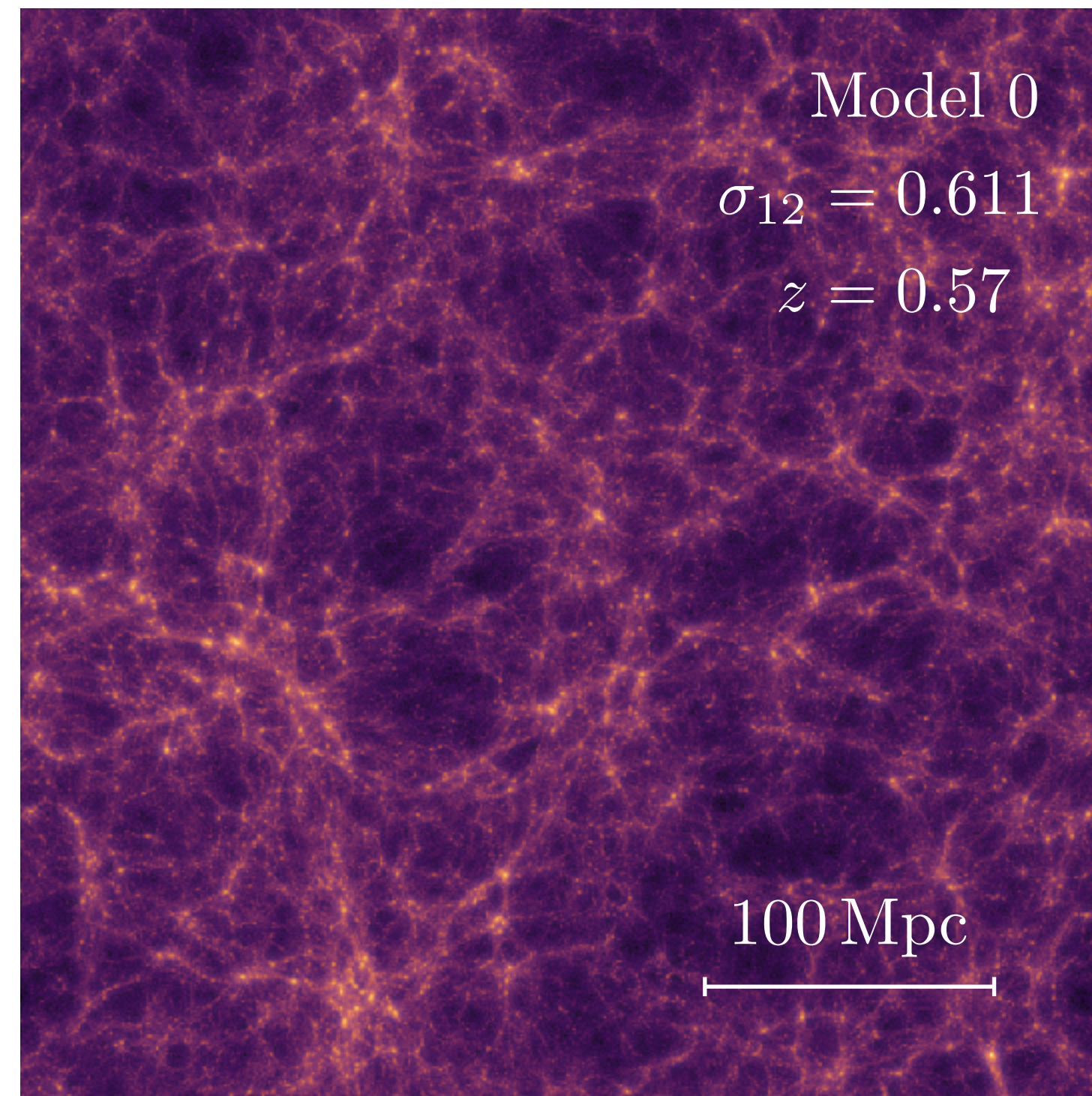
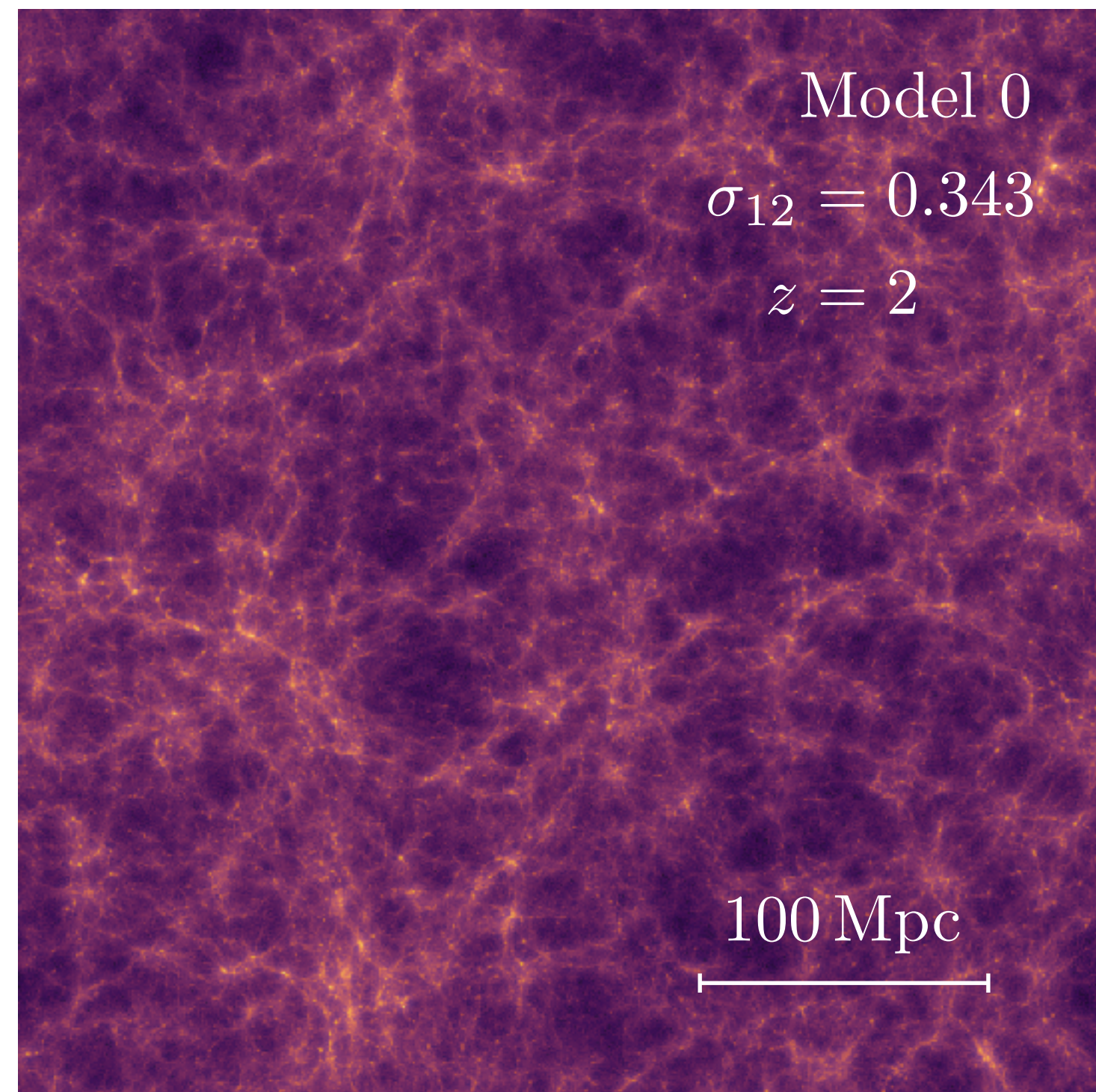
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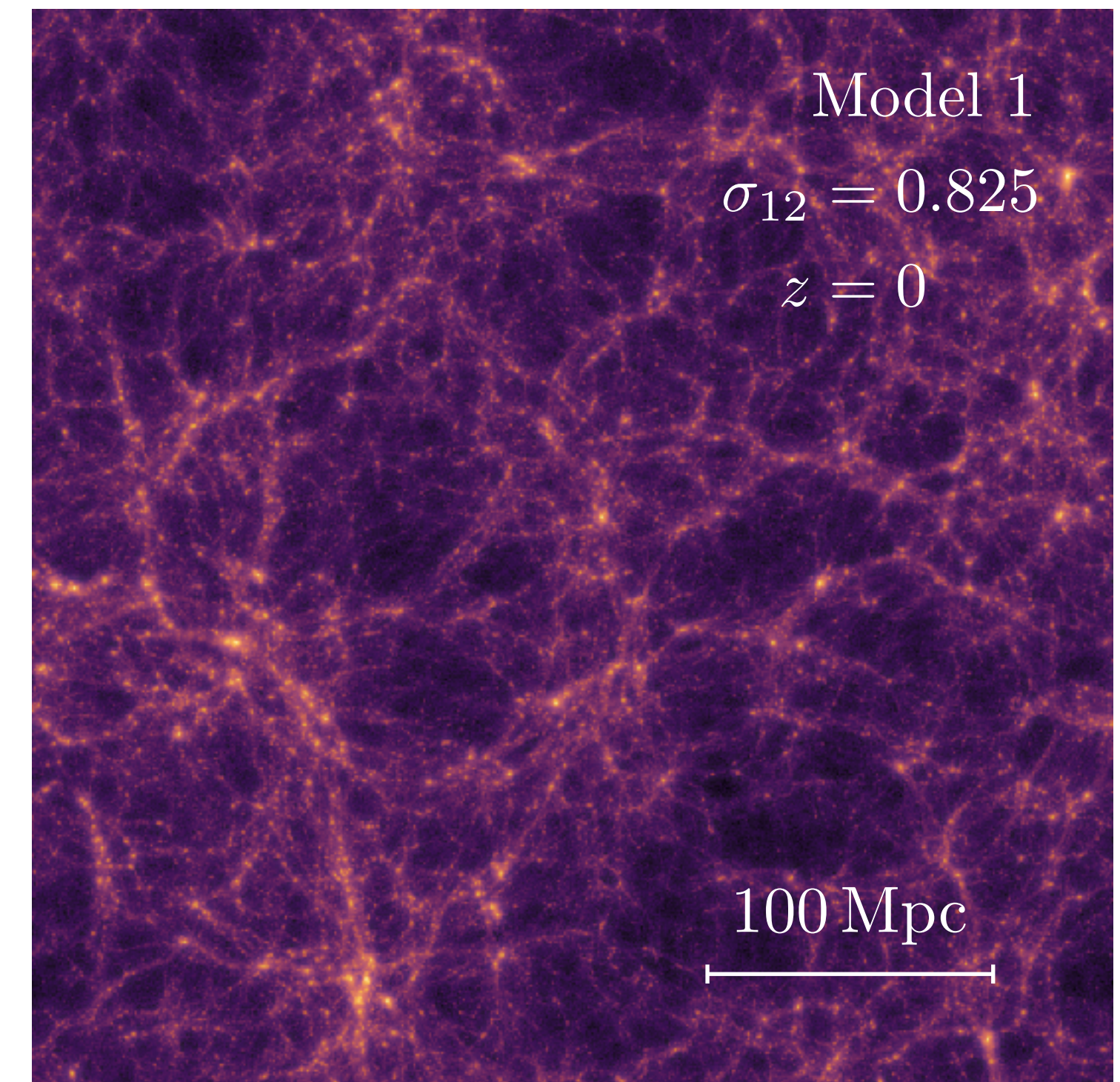
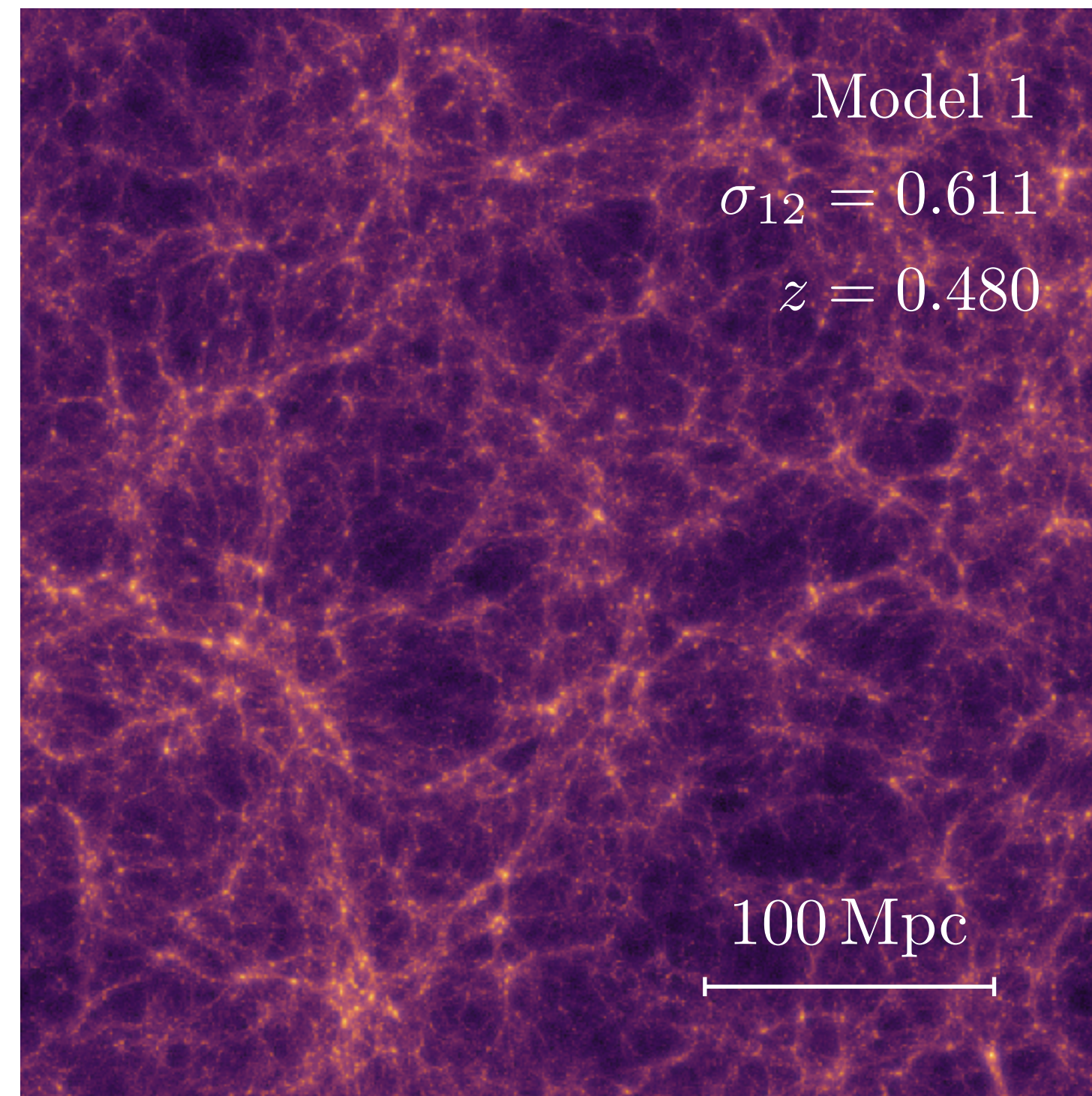
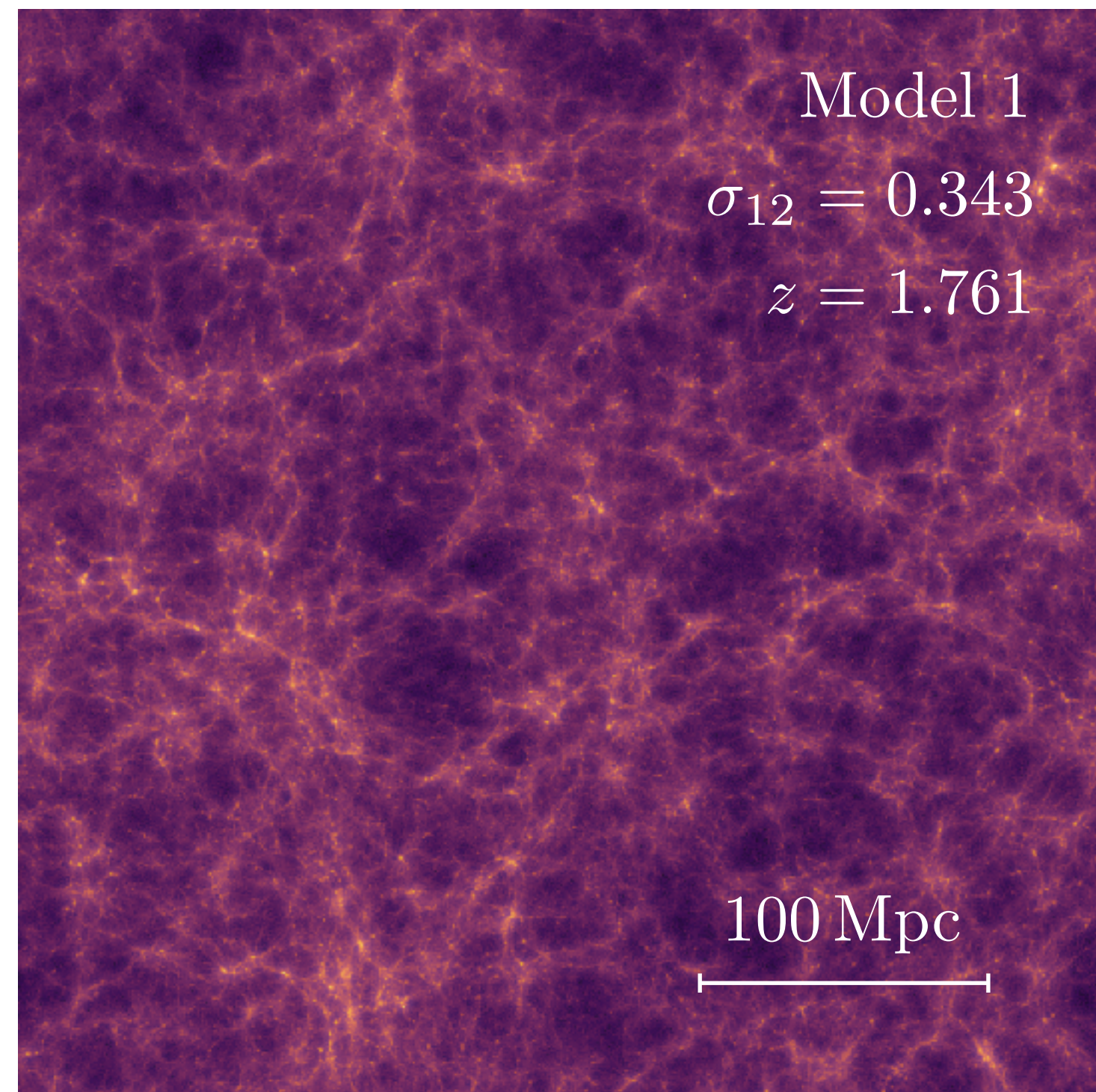
# Beyond two-point statistics

- Evolution mapping describes the full density field.
- It can be used to describe multiple statistics.



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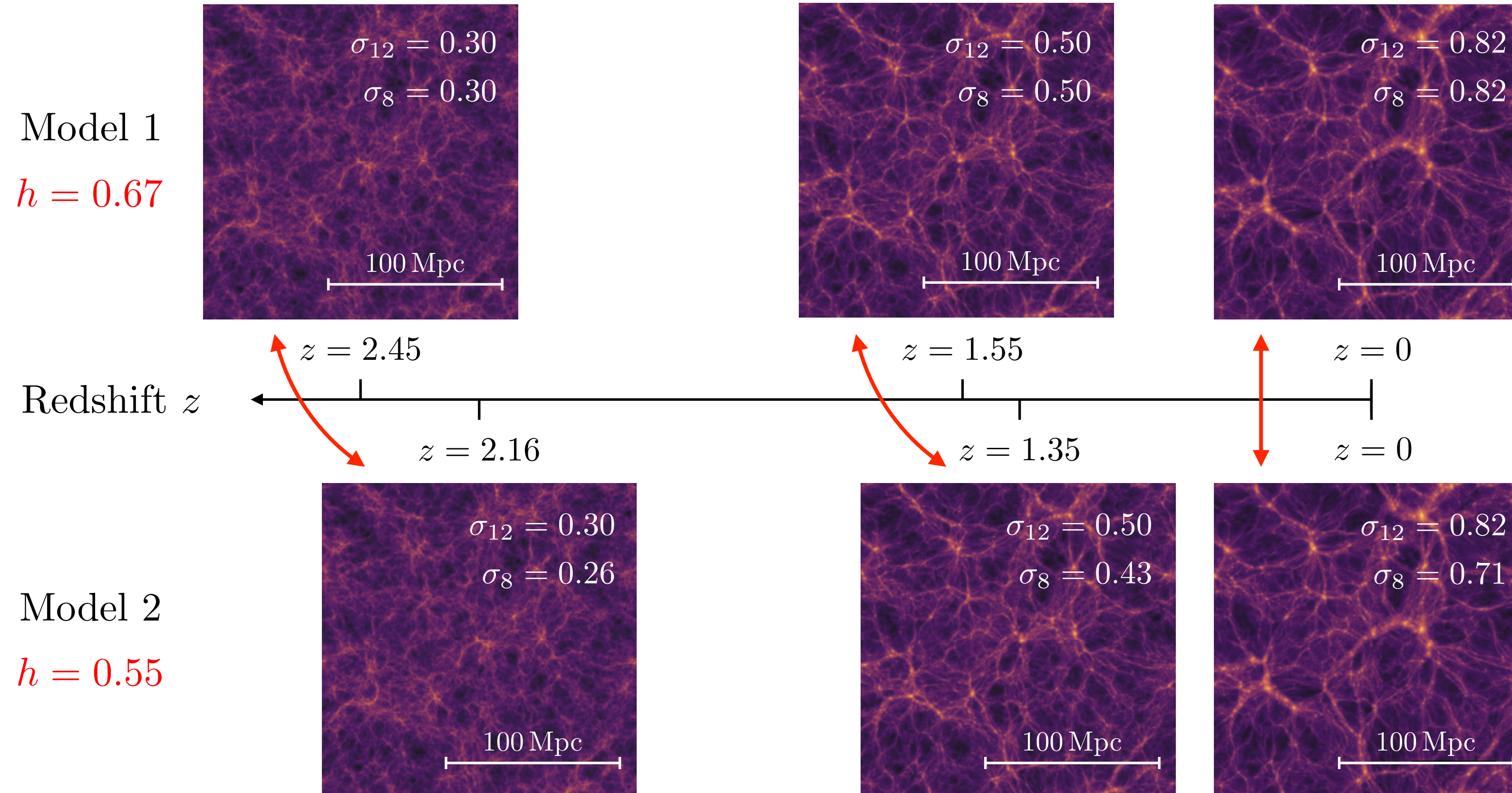


# Emulating the non-linear $P(k)$

Standard approach:  $\Theta = (\underbrace{\omega_c, \omega_b, n_s}_{\text{shape}}, \underbrace{\omega_K, \omega_{\text{DE}}, w_0, w_a, \dots}_{\text{evolution}}, z)$

Evolution mapping:  $\Theta = (\omega_c, \omega_b, n_s, \sigma_{12})$

## ALETHEIA



Evolution mapping reduces the required number of parameters to describe  $P(k|z)$ .

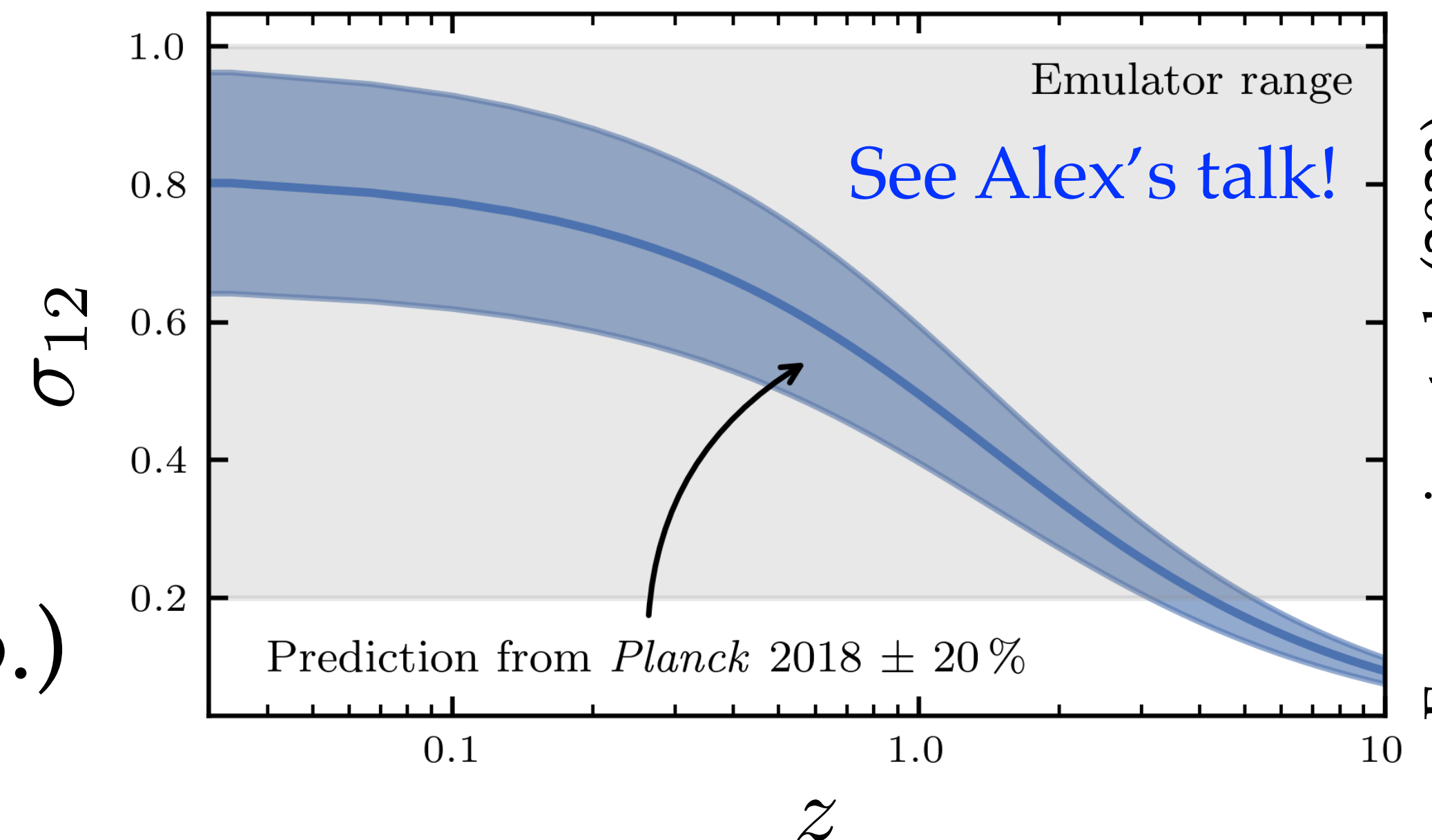
Emulator results must be corrected by  $\Delta g(\sigma_{12})$

# COMET: Emulating perturbation theory

- Evaluation of  $P_\ell(k|z)$  takes a few seconds  
-> MCMC analyses require a few days.
- For PT models, evolution mapping is exact.
- For a reference set  $\Theta_{e,0}$ , we sample
$$\Phi = (\Theta_s, \sigma_{12}, f)$$
- COMET is available as a Python package  
<https://pypi.org/project/comet-emu/>
- New versions adding  $\omega_\nu$  (Pezzotta+ in prep.)  
and config. space (Semenait+ in prep.)

COMET - Cosmological Observables  
Modelled by Emulated perturbation Theory

Parameter	Min. emulator range	Max. emulator range
$\omega_b$	0.0205	0.02415
$\omega_c$	0.085	0.155
$n_s$	0.92	1.01
$\sigma_{12}$	0.2	1.0
$f$	0.5	1.05

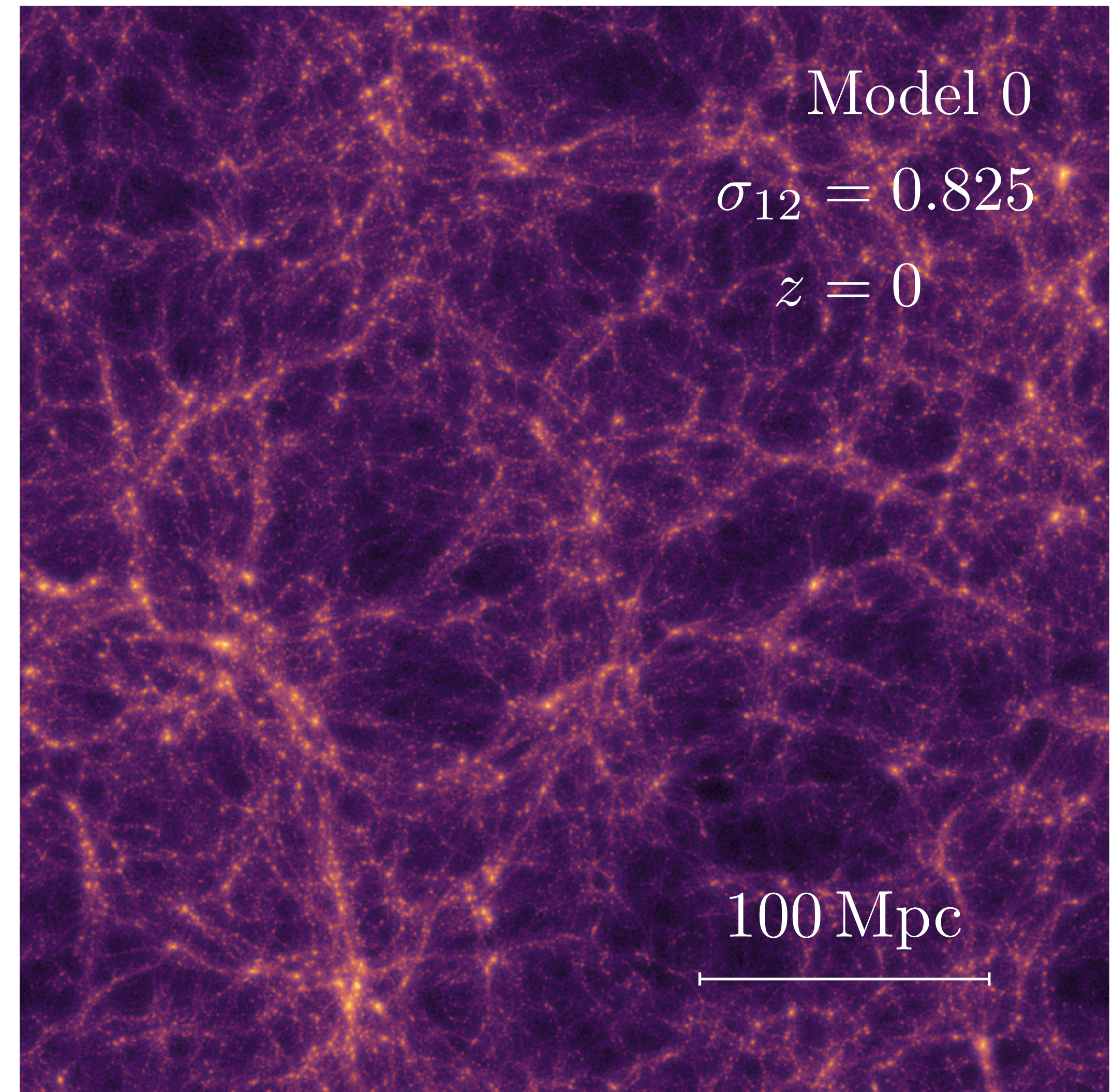


# Field-level emulation

The computational cost of N-body simulations hinders the use of SBI.

CNNs can reproduce full N-body simulations based on their linear inputs (e.g., He et al. 2018, Jamieson et al. 2023).

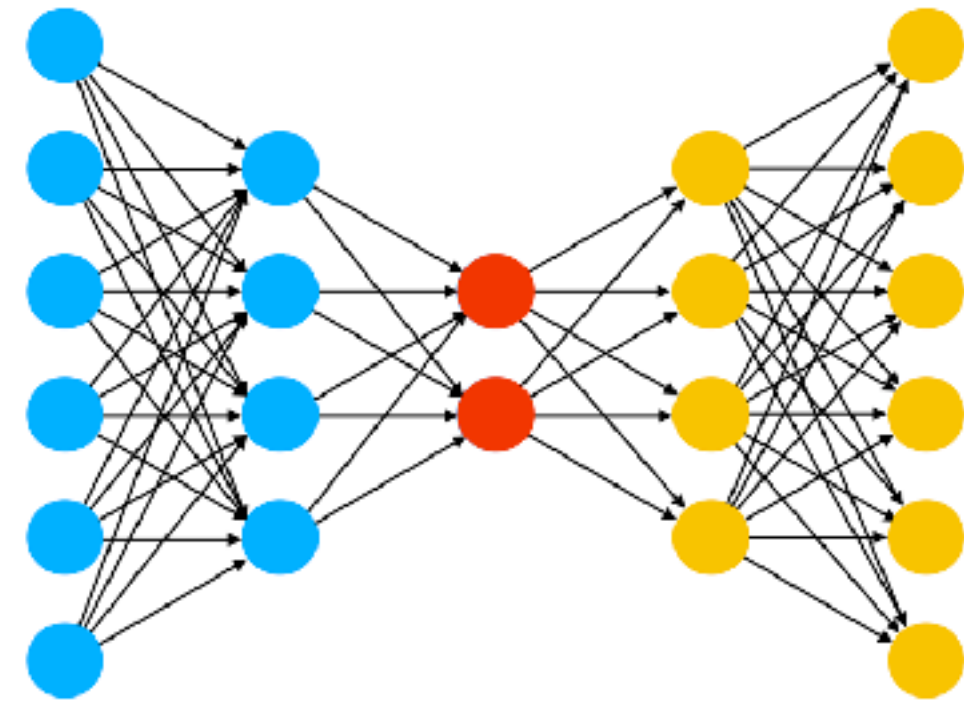
Evolution mapping can help to generalise these results.



# Field-level emulation

**Input:** 2LPT  
particle positions

$$\mathbf{x}_{\text{in}} = \mathbf{x}_{2\text{LPT}}$$



**Output:** shift to N-body  
particle position

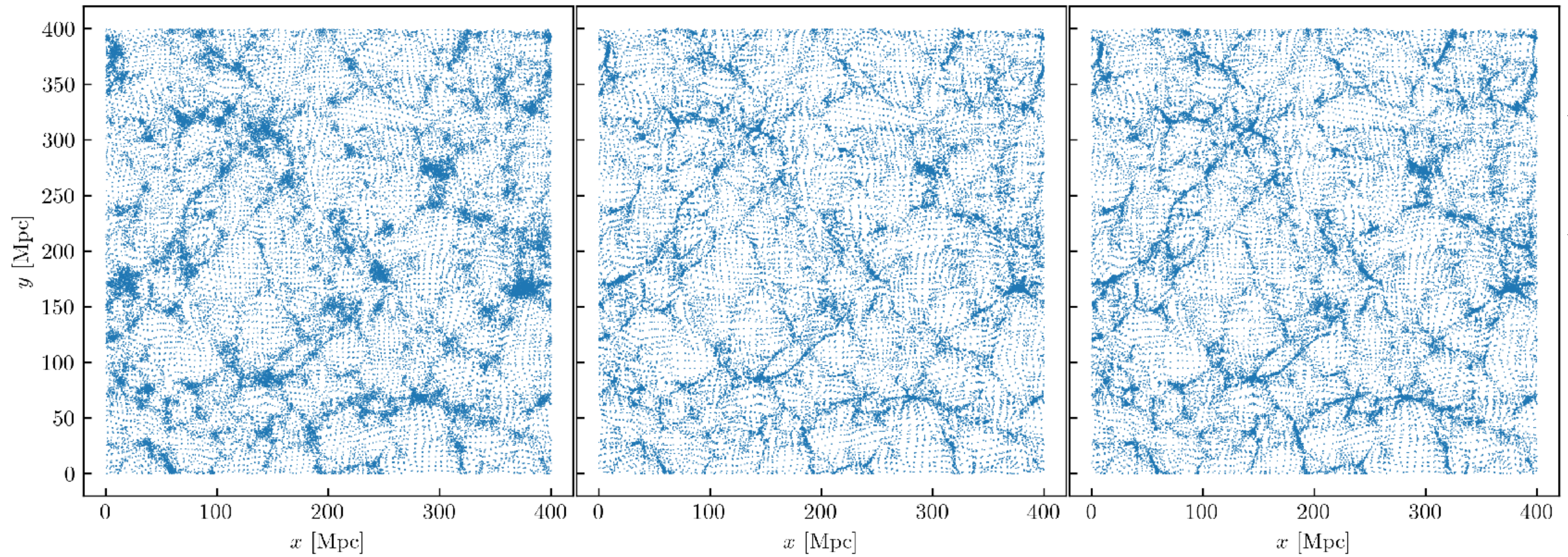
$$\mathbf{x}_{\text{out}} = \mathbf{x}_{\text{Nbody}} - \mathbf{x}_{2\text{LPT}}$$

2LPT

N-body

Emulator

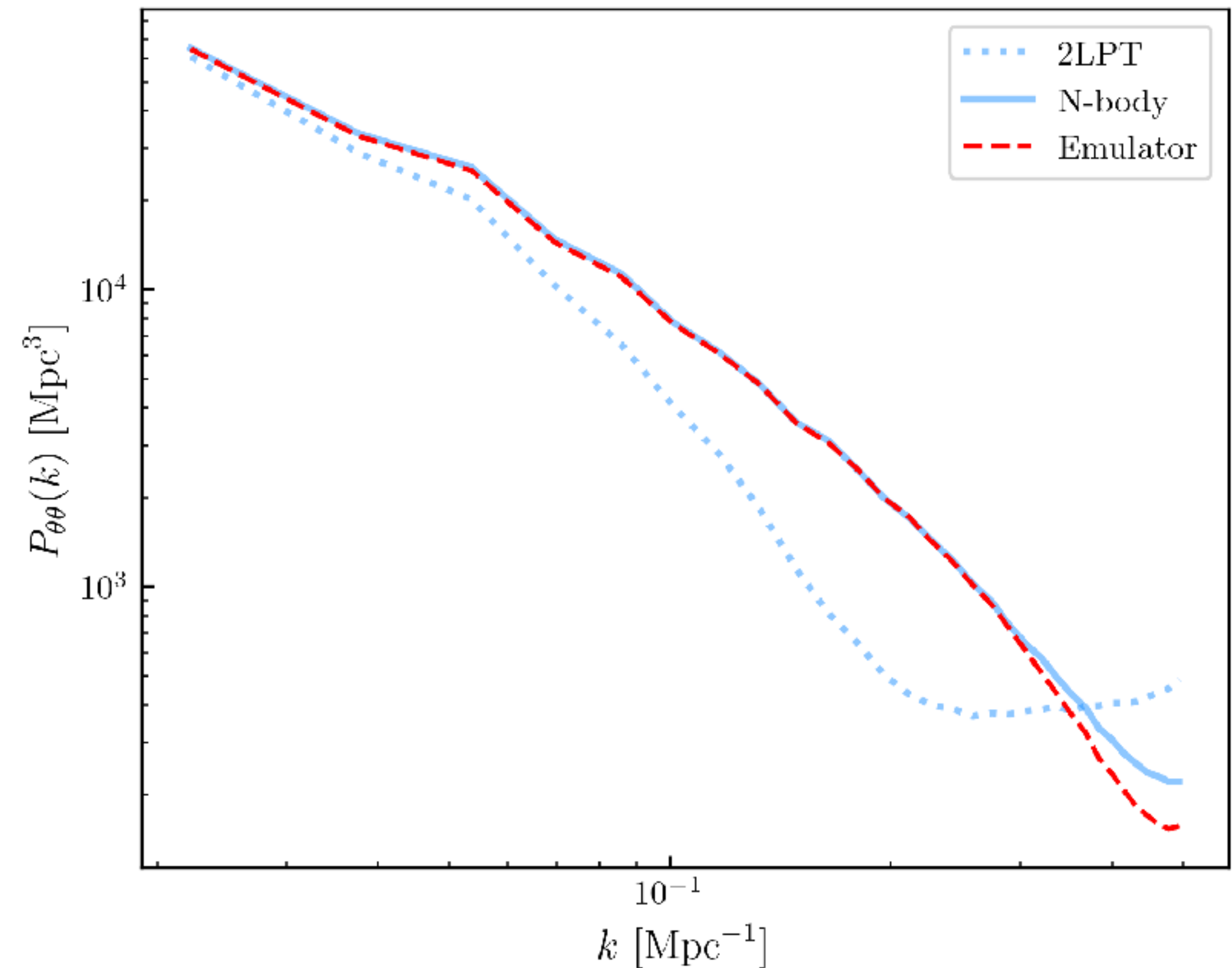
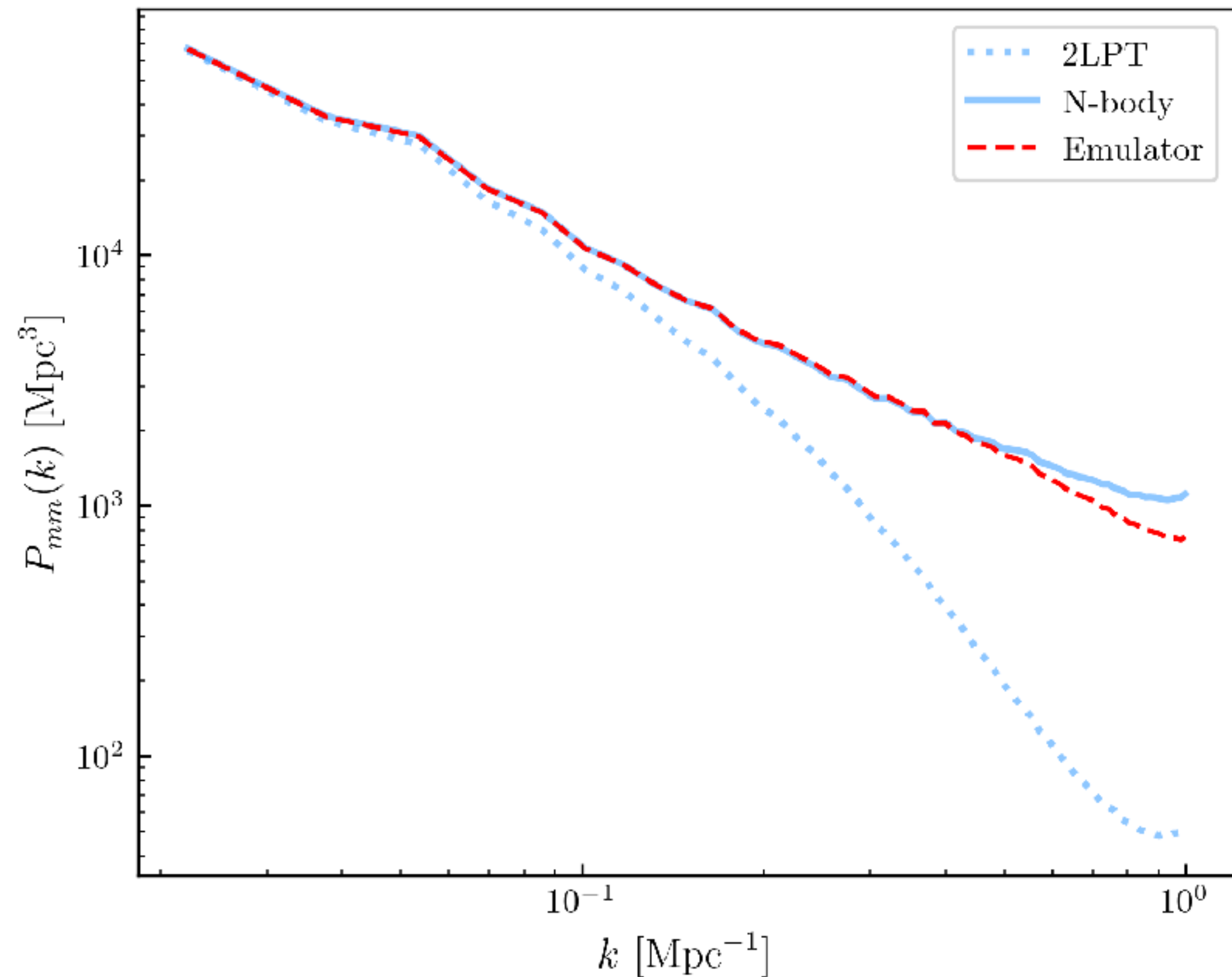
Preliminary!



Correa et al. (in prep.),  
Perez Fernandez et al. (in prep.)

# Field-level emulation

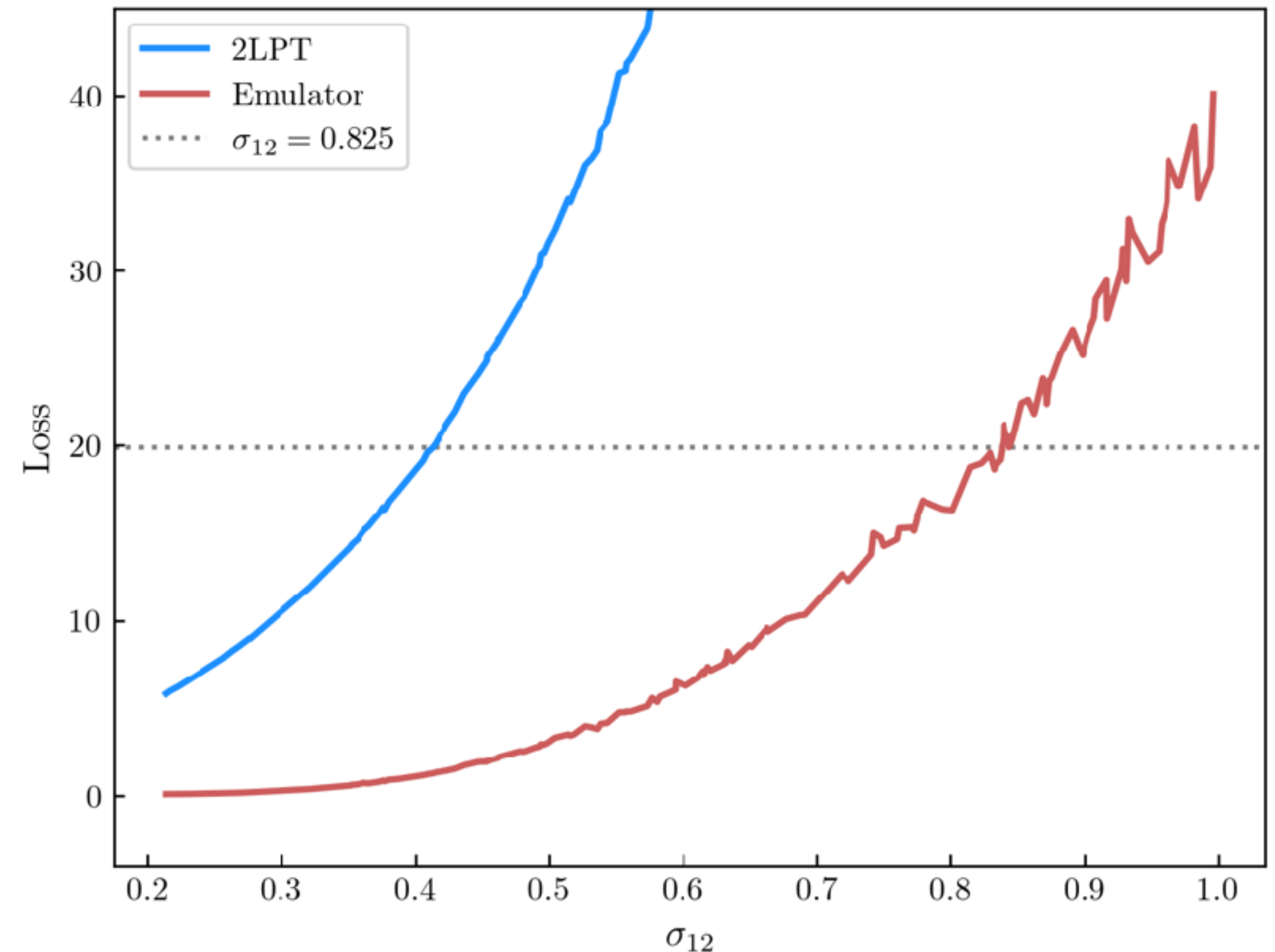
- Preliminary results show good performance.
- Currently studying the cosmological dependency of the results



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# Field-level emulation

- Preliminary results show good performance.
- Currently studying the cosmological dependency of the results
- Main parameter controlling the emulator's performance is  $\sigma_{12}(z)$ .
- Testing the impact of different structure formation histories.
- Apply to extensions of  $\Lambda$ CDM.



Correa et al. (in prep.),  
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# The information content of $P(k, \mu)$

- Parameter degeneracies are modified for biased tracers in redshift-space

$$P_{\text{gg}}(k, \mu) = (b_1 \sigma_{12} + f \sigma_{12} \mu^2)^2 \frac{P_{\text{mm}}(k)}{\sigma_{12}^2}$$

- For fixed  $\Theta_s$ , models with the same values of  $b_1 \sigma_{12}$  and  $f \sigma_{12}$  are identical.
- The BAO signal provides constraints on

$$D_{\text{M}}(z)/r_{\text{d}}$$

$$D_{\text{H}}(z)/r_{\text{d}}$$

- The broad-band shape of  $P_{\text{gg}}(k, \mu)$  contains weak information on the shape parameters (e.g.,  $n_s$ ).

# LSS analysis methods

## “Full-modelling” approach:

- Select parameter space to be explored: e.g.,  $\Lambda$ CDM,  $\Theta = (\omega_b, \omega_c, \omega_{DE}, n_s, A_s)$
- Theoretical predictions directly compared against clustering data.

## “Template” approach:

- Assume a fixed template cosmology.
- Differences between data and the template are compressed into:

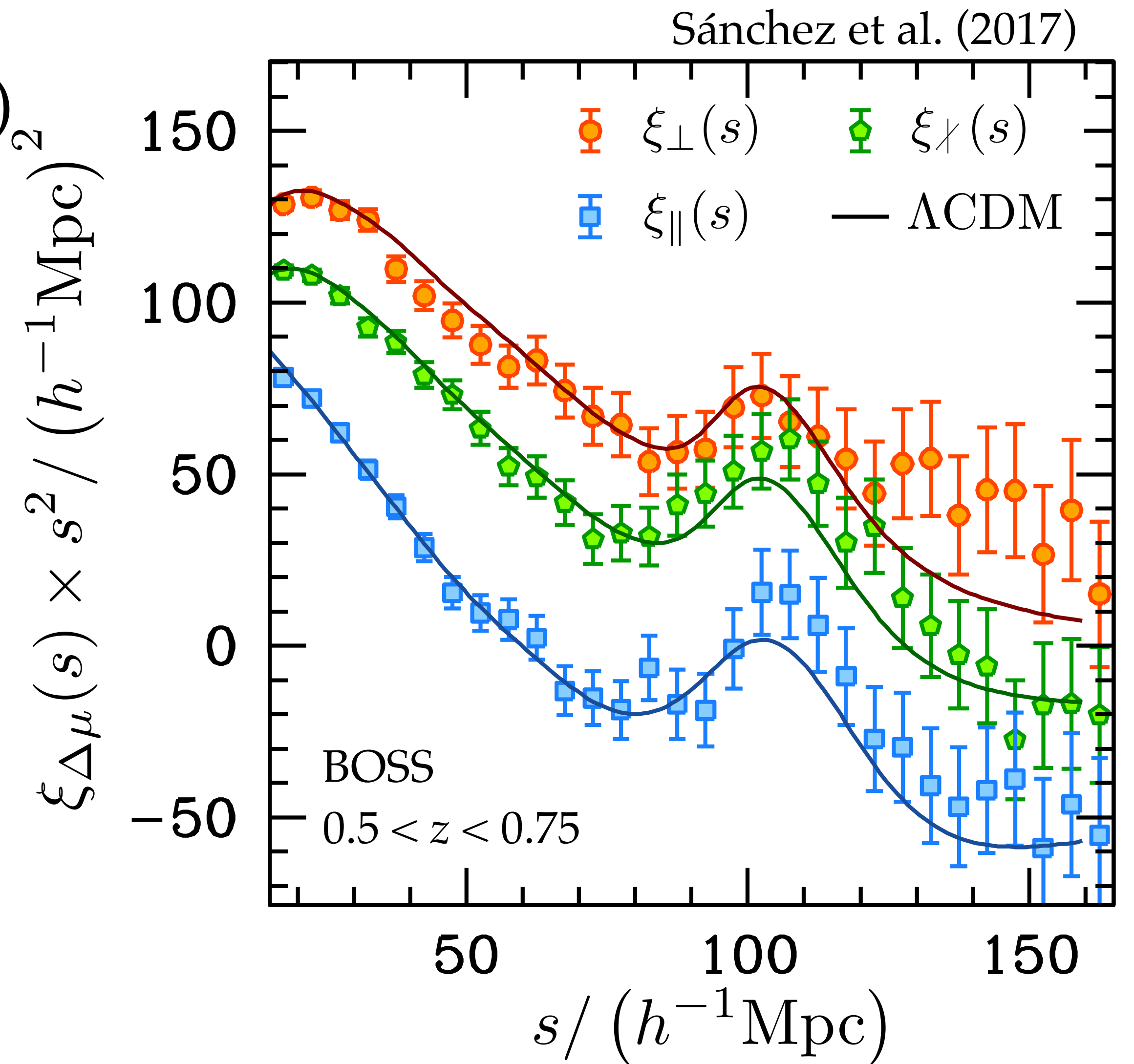
$$D_M(z)/r_d, D_H(z)/r_d, f\sigma_{8/h}(z)$$

- *Shapefit* includes two parameters,  $m, n$ , describing the shape of  $P(k)$



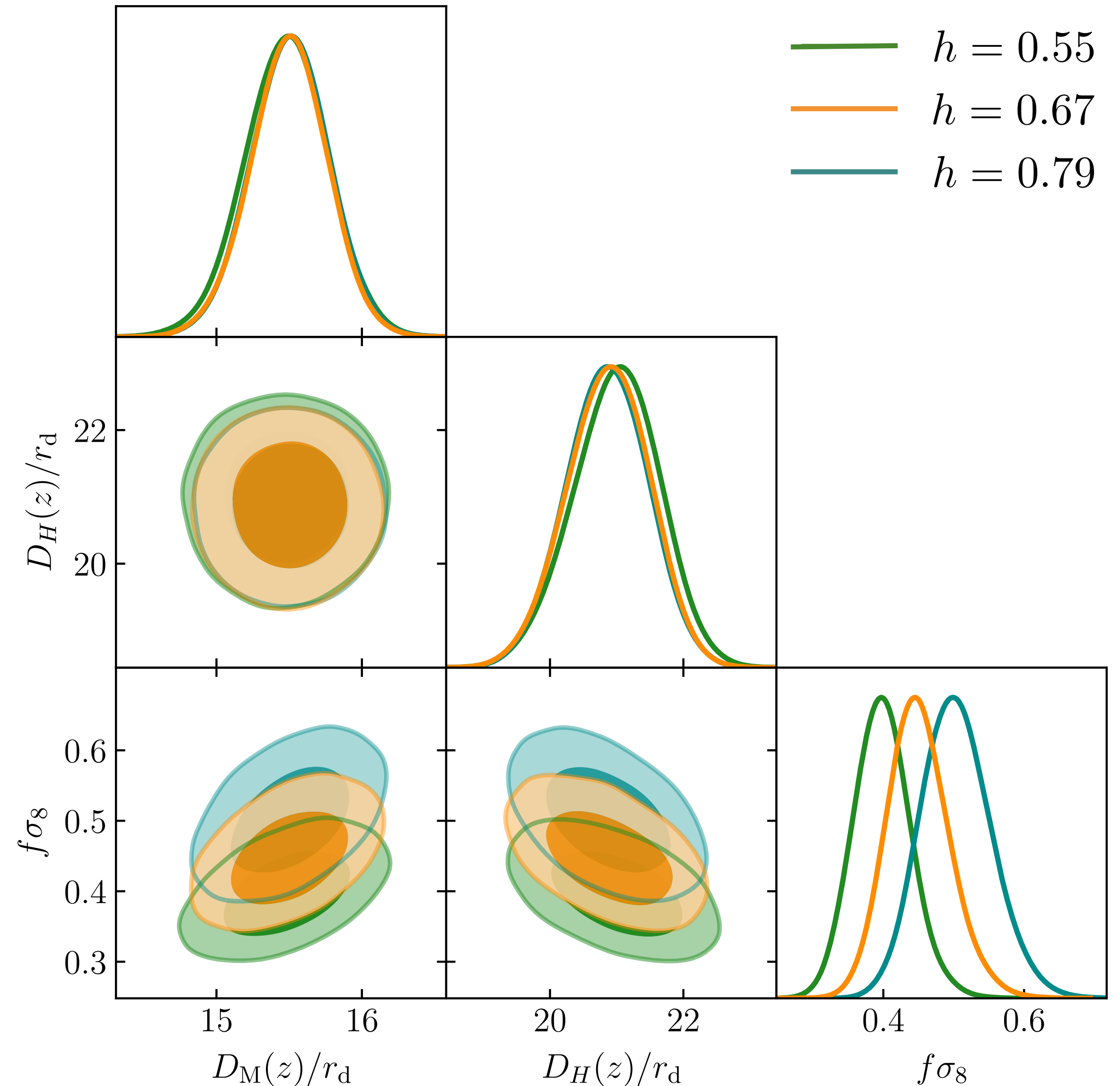
# Full-modelling LSS analyses

- Most LSS studies used full modelling (Sánchez+ 2017, Semenaite+ 2022, 2023)
- Focus on accuracy of the constraints: analyses used LSS + CMB data.
- Current focus: asses the consistency between different data sets.
- Several BOSS-only analyses (d'Amico+ 2020, Ivanov+ 2020, Tröster + 2020, ...)



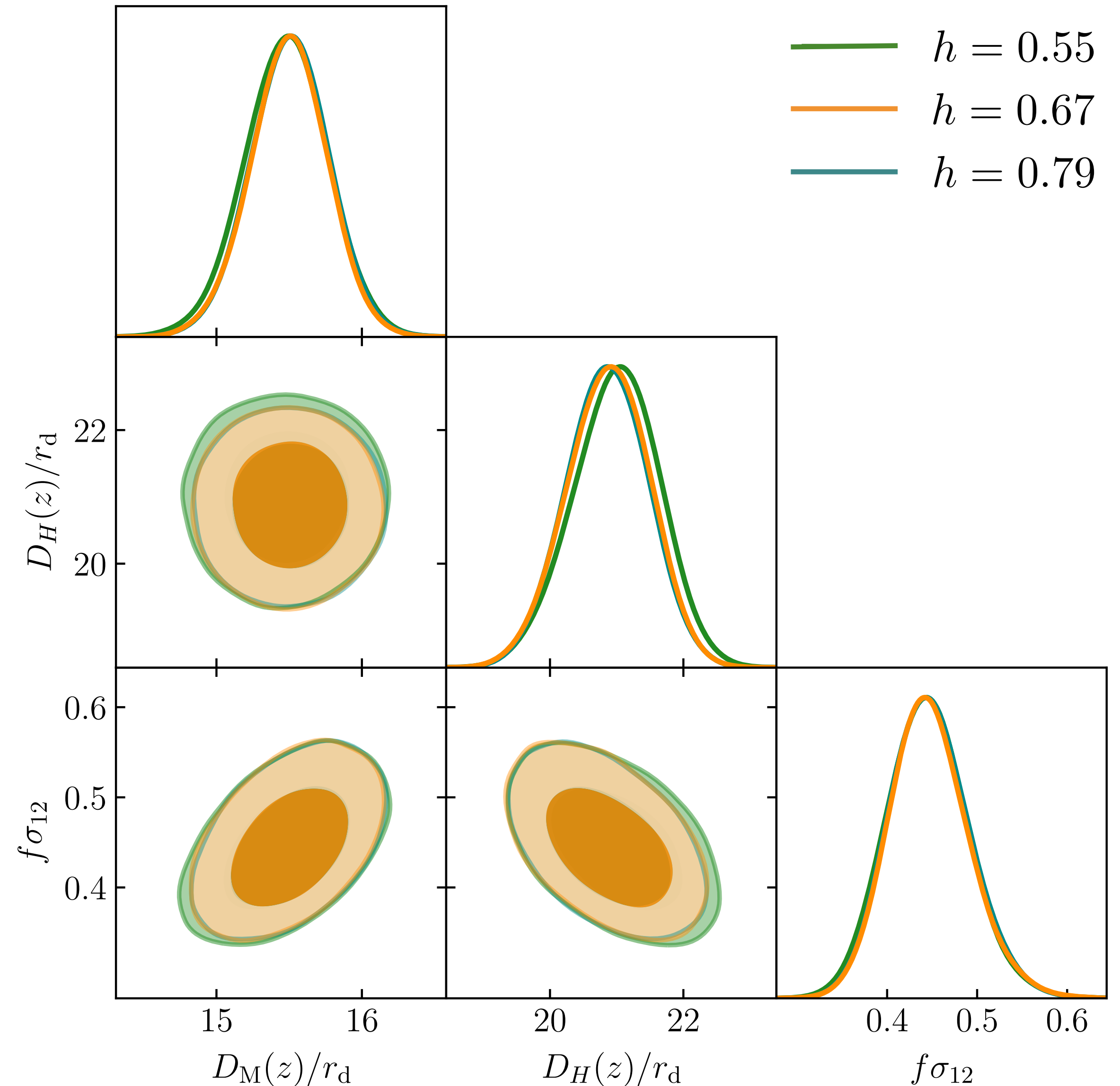
# Template LSS analyses

- In the standard template analysis,  $h$  is kept fixed.
- The constraints on  $f\sigma_{8/h}(z)$  depend on that assumption.
- The correct error on  $f\sigma_{8/h}(z)$  should be marginalised over  $h$ .
- The effect disappears when expressed in terms of  $f\sigma_{12}(z)$ .



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# Template LSS analyses

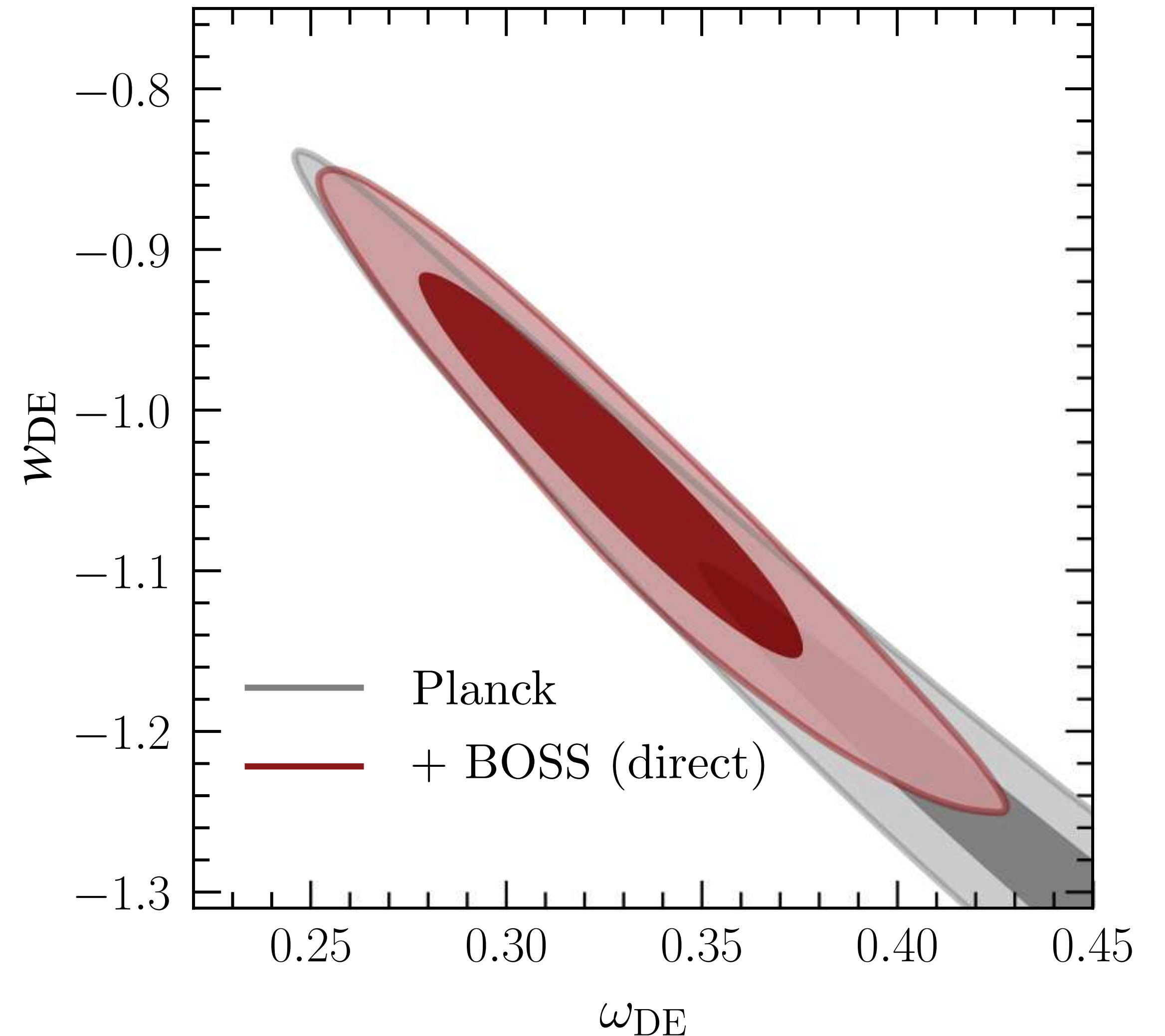
- Constraints on  $w$ CDM:

$$\Theta = (\omega_b, \omega_c, \omega_{\text{DE}}, n_s, A_s, w_{\text{DE}})$$

- Planck + BOSS (high-z)

- *Full-modelling* analysis:

$$w_{\text{DE}} = -1.04 \pm 0.082$$



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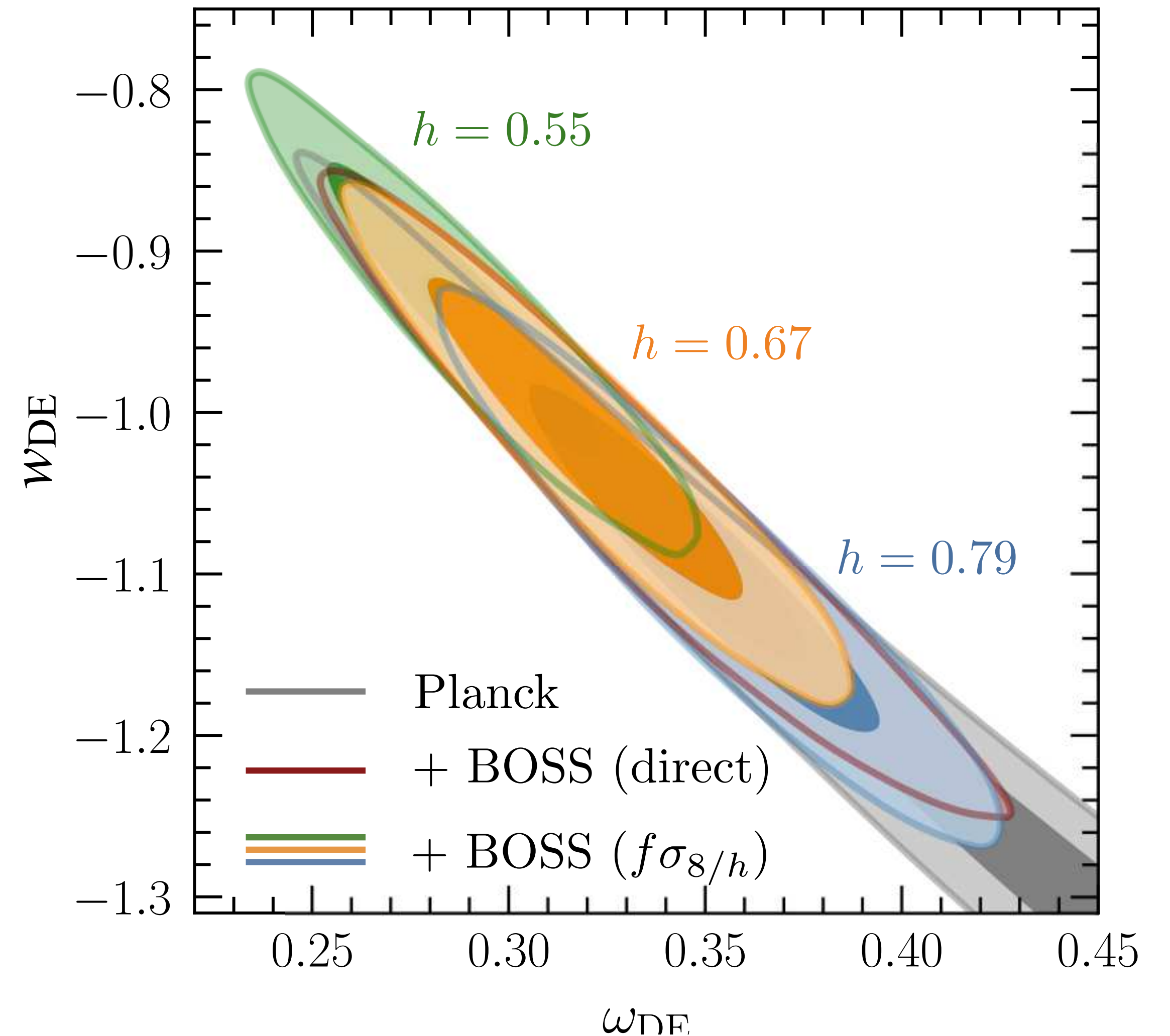
- Planck + BOSS (high- $z$ )

- *Full-modelling* analysis:

$$w_{\text{DE}} = -1.04 \pm 0.082$$

- *Template* analysis using  $f\sigma_{8/h}(z)$ :

$$w_{\text{DE}} = -1.02 \pm 0.065$$



# Template LSS analyses

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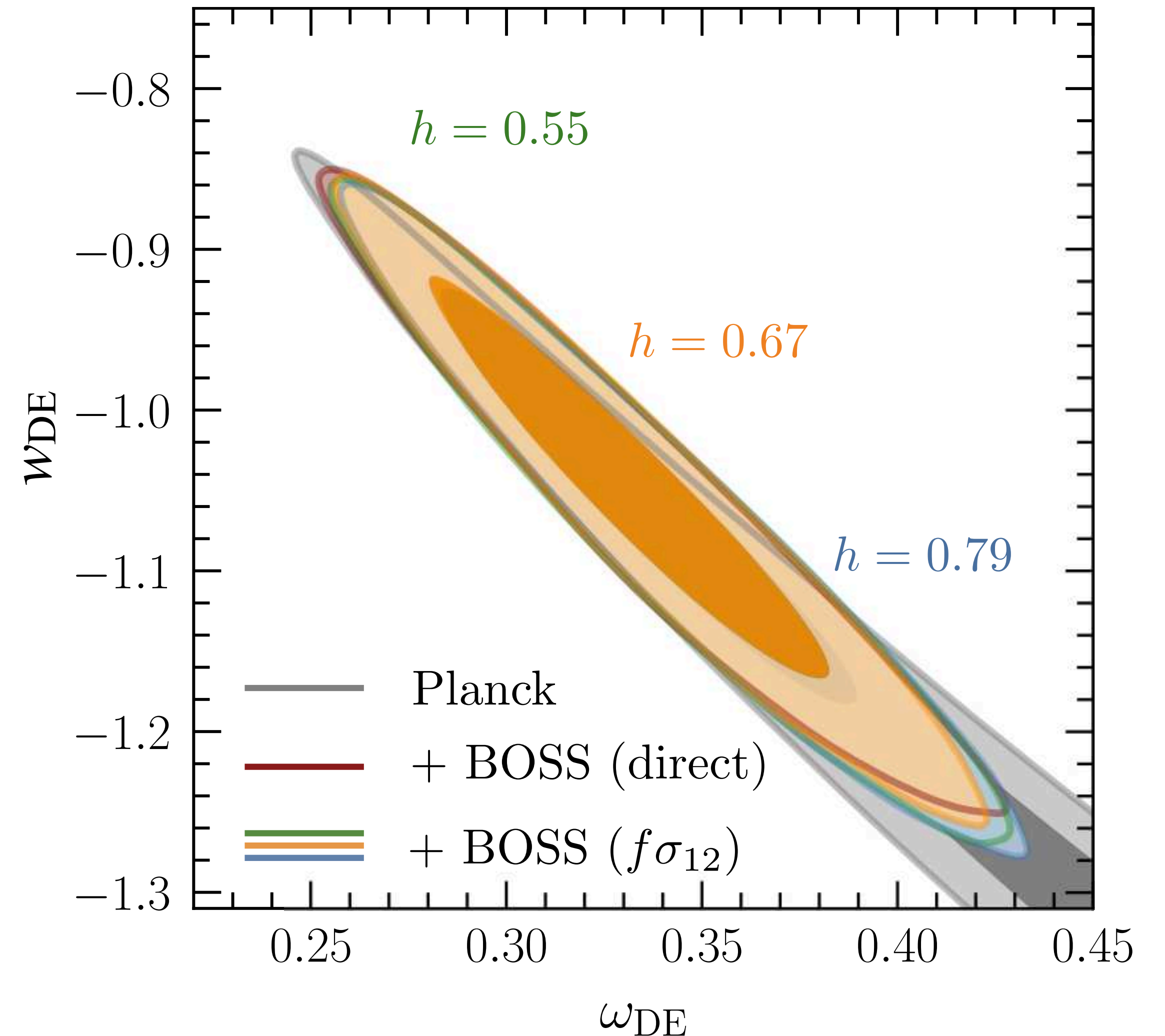
$$w_{\text{DE}} = -1.04 \pm 0.082$$

- *Template* analysis using  $f\sigma_{8/h}(z)$ :

$$w_{\text{DE}} = -1.02 \pm 0.065$$

- *Template* analysis using  $f\sigma_{12}(z)$ :

$$w_{\text{DE}} = -1.05 \pm 0.083$$



# Template LSS analyses

- Gil-Marin+ (2020) proposed to use  $f\sigma_{8,q}$

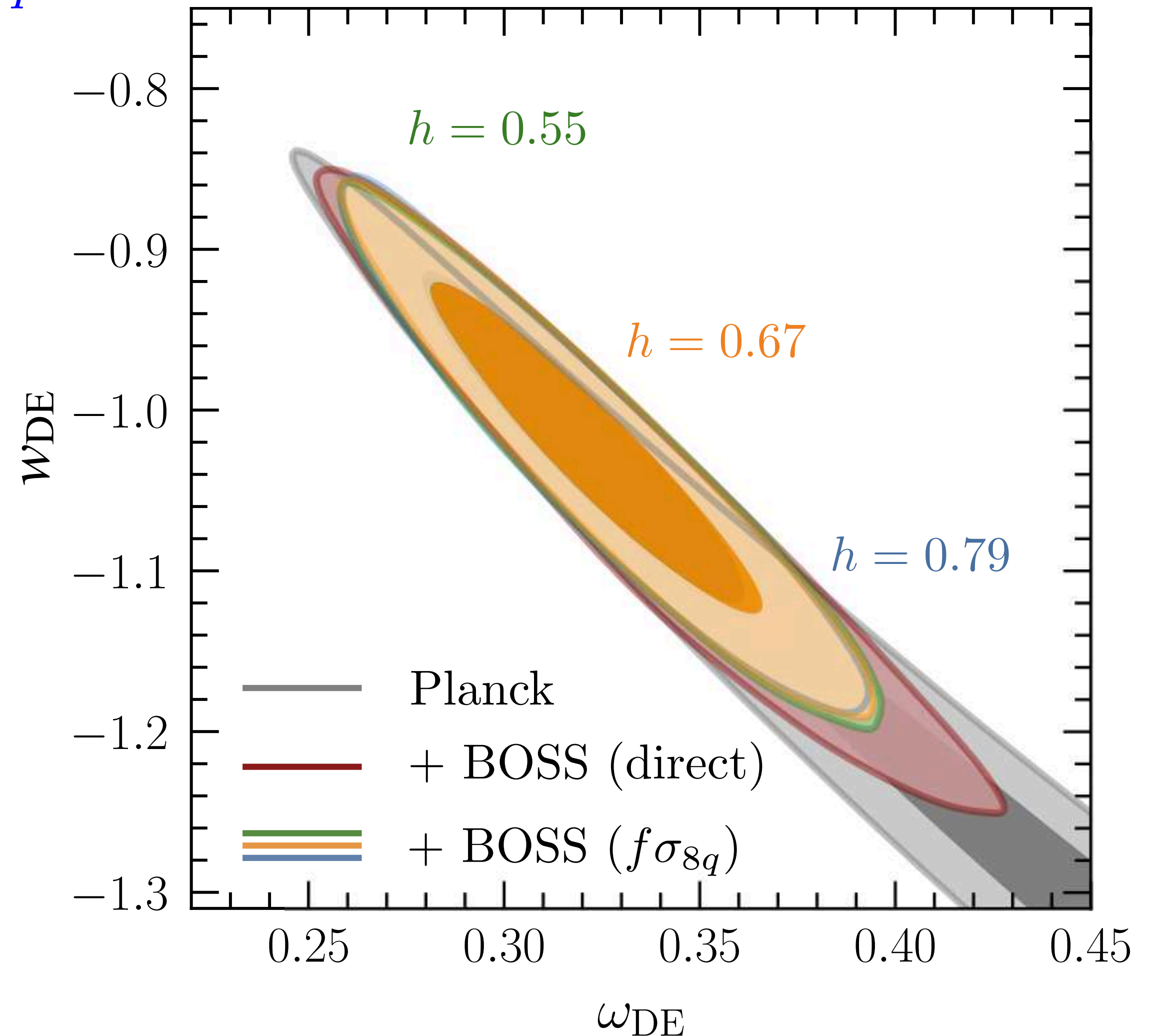
$$\sigma_{8,q}^2 = \int_0^\infty dk k^2 P_L(k) W^2(s_8 q k),$$

where  $s_8 = (8/h)$  Mpc and  $q^3 = q_\perp^2 q_\parallel$

- This quantity cannot be used as the standard  $f\sigma_{8/h}(z)$ .

- Interpreting  $f\sigma_{8,q}$  as  $f\sigma_{8/h}$  leads to

$$w_{\text{DE}} = -1.03 \pm 0.065$$

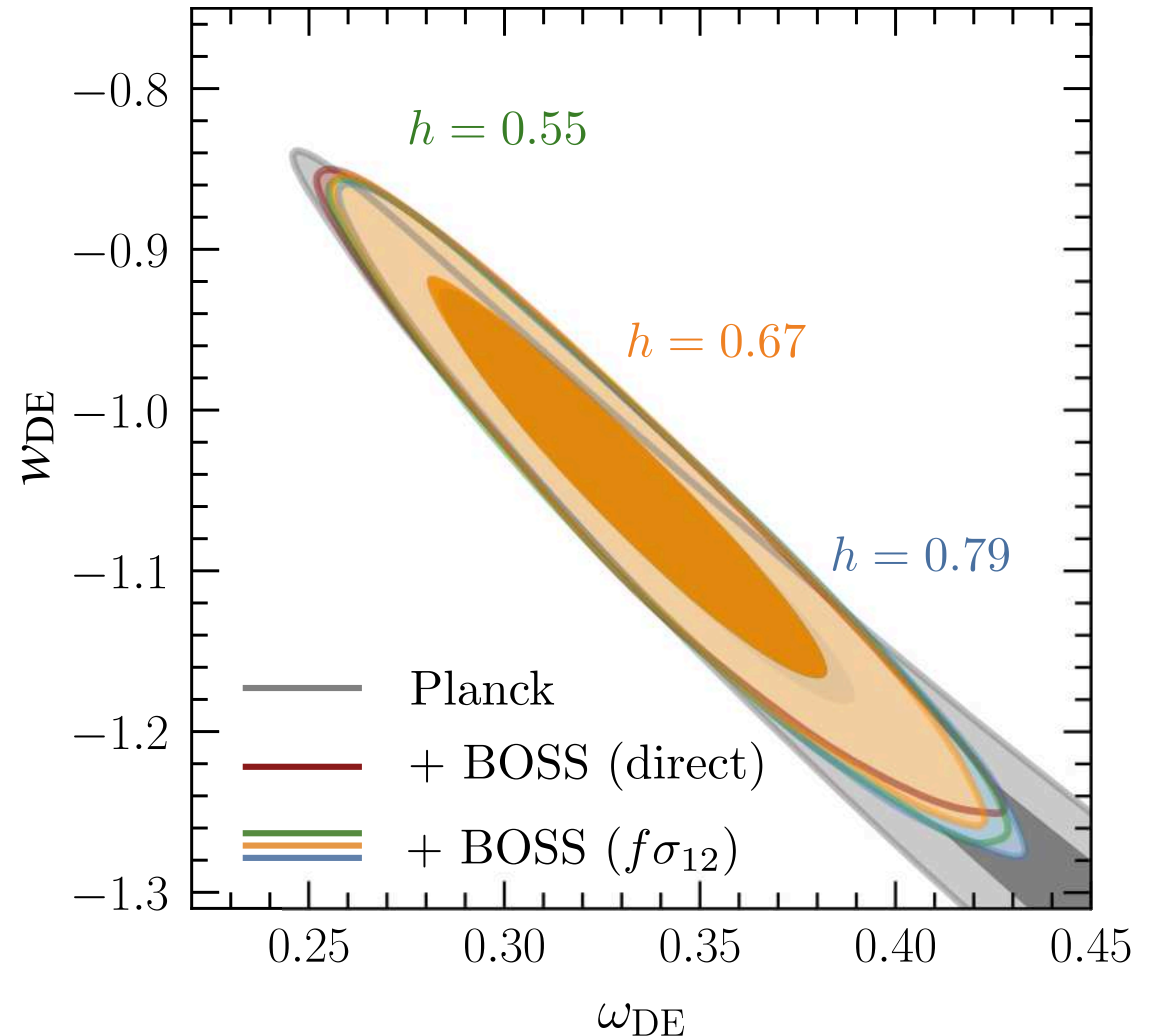


# Template LSS analyses

- Brieden+ (2021, SHAPEFIT) proposed to use instead

$$s_d = \frac{r_d}{r_{d,\text{fid}}} (8/h) \text{ Mpc}$$

- A scale proportional to  $r_d$  means using a reference scale in **Mpc**.
- Multiple analyses still quote their results in terms of  $f\sigma_{8/h}(z)$ .





# Final remarks

- Evolution mapping: we classify parameters into *Shape* and *evolution* based on their impact on  $P_L(k|z)$ .
- At the linear level,  $\Theta_e$  follow a perfect degeneracy, described by  $\sigma_{12}$ .
- This is partially inherited by the non-linear density field, with deviations sensitive to the suppression  $g(a) = D(a)/a$ .
- We are using evolution mapping to build new descriptions of the non-linear matter density field.
- This approach can help us to better understand the information content of all clustering measurements.